

Neutrino mass measurement with Ho

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on behalf of HOLMES collaboration*

HOLMES collaboration:
INFN Ge, INFN Mib, INFN LNGS, Unimib, NIST, Lisbona, PSI, ILL

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- ^{163}Ho Electron capture and Neutrino mass measurement
- Neutrino mass Experiment with ^{163}Ho
 - ECHo
 - HOLMES
 - LALN
- Conclusion



Kinematics of weak decays

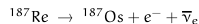
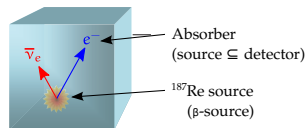
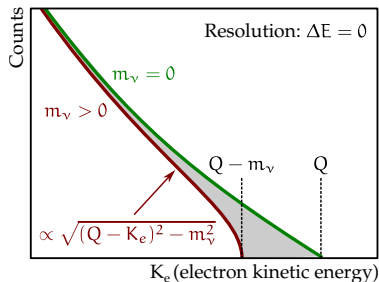
- nuclear decay involving neutrinos
- only energy and momentum conservation
- no further assumptions

General experiment requirements:

- high statistics at the beta spectrum end-point
 - low E_0 β decaying isotopes: ^{187}Re , ^{163}Ho
- high energy resolution ΔE

Calorimeters: source $\subseteq$ detector:

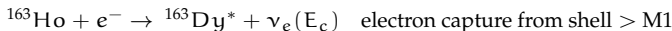
- the β source is embedded in the detector
- ideally they measure all the energy E released in the decay except for the ν_e energy
- calorimeters measure the entire spectrum at once



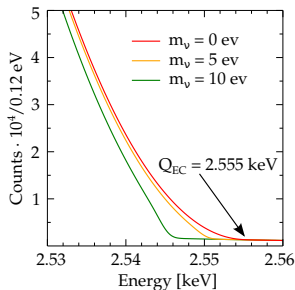
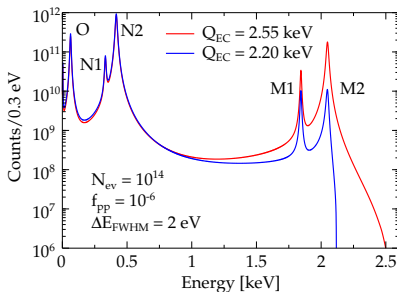
The electron capture of the ^{163}Ho



An interesting isotope suitable for the neutrino mass experiment could be the ^{163}Ho .



proposed by A. De Rujula e M. Lusignoli in 1982 [1, 2]



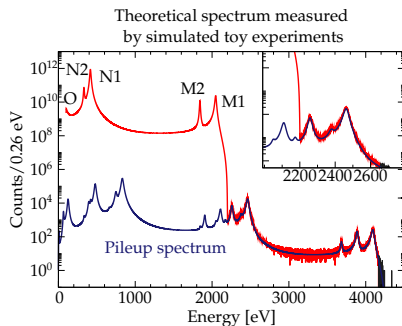
$$\frac{d\lambda_{EC}}{dE_c} = \frac{G_\beta^2}{4\pi^2} (Q_{EC} - E_c) \sqrt{(Q_{EC} - E_c)^2 - m_\nu^2} \times \sum_i n_i C_i \beta_i^2 B_i \frac{\Gamma_i}{2\pi} \frac{1}{(E_c - E_i)^2 + \Gamma_i^2/4},$$

- Series of lines at the ionization energies E_i of the captured electrons
- The peaks have Breit-Wigner shapes (width Γ_i of a few eV)
- End-point shaped by $\sqrt{(Q - E_e)^2 - m_\nu^2}$ (the same of the β -decay)
- Self calibrating spectrum



- Calorimetric measurement of Dy atomic de-excitations (mostly non-radiative):
 - ⇒ measurement of the entire energy released except the ν energy
- The rate at end-point may be as high as for ^{187}Re but depends on Q_{EC} :
- Q_{EC} and atomic de-excitation spectrum poorly known:
 - ⇒ Measured: $Q_{\text{EC}} = (2.2 \div 2.8) \text{ keV}$
 - ⇒ Recommended: $Q_{\text{EC}} = 2.555 \text{ keV}$ [3, 4])
- $\tau_{1/2} \simeq 4570 \text{ years} \Rightarrow$ high specific activity:
 - ⇒ Holmium detector not needed
 - ⇒ ^{163}Ho can be implanted in any suitable microcalorimeter absorber
- Complex pile-up spectrum
- No high statistics and clean calorimetric measurement so far

^{163}Ho : pile-up spectra



$$S(E_c) = [N_{ev}(N_{EC}(E_c, m_v) + f_{pp} \times N_{EC}(E_c, 0) \otimes N_{EC}(E_c, 0)) + B(E_c)] \otimes R_{\Delta E}(E_c)$$

N_{ev} : total number of events
 $N_{EC}(E_c, m_v)$: ^{163}Ho spectrum
 $B(E)$: background energy spectrum
 $R_{\Delta E}(E_c)$: detector energy response function
 f_{pp} : fraction of pile-up events
 $R_{\Delta E}(E_c)$: detector energy response function
 ΔE : interval of energy

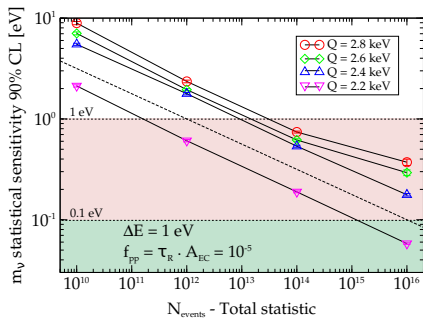
more details in A. Nucciotti, submitted to EPJC,
arXiv:1405.5060 [5]

- Pulse pile-up occurs when multiple events arrive within the temporal resolving time of the detector
- Unresolved pile-up produces background close to the end-point
- The ^{163}Ho pile-up events spectrum is quite complex and presents a number of peaks right at the end-point of the decay spectrum
- To resolve pile-up:
 - Detector with fast signal rise-time τ_{rise}
 - Pile-up recognition algorithm



Montecarlo Simulation

- The approach is a replica of the one outlined in [6] for beta decay calorimetric experiments:
 - large number of simulated spectra that would be measured by a large number of experiments carried out in a given configuration
 - the spectra are analyzed as the real ones
 - the statistical sensitivity is deduced from the distribution of the obtained m_ν^2



A. Nucciotti, submitted to EPJC, arXiv:1405.5060 [5]

Requirements:

- High energy resolution ($\simeq 1$ eV)
- Fast response detectors ($\simeq 1$ μs) to avoid pile-up events
- Multiplexable detectors array ($\simeq 1000$)



TESs, MMCs, MKIDs, ...

Holmium LTD experiment status



^{163}Ho seems to be better than ^{187}Re :

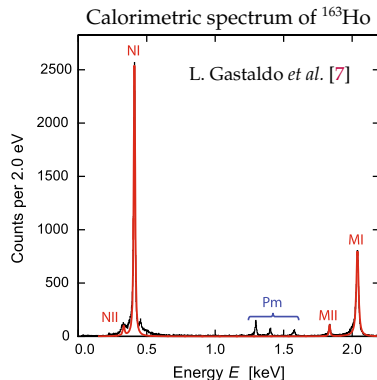
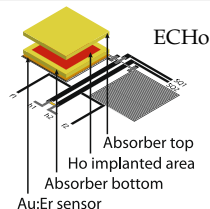
- higher specific activity $\Rightarrow$ Holmium detector not needed
- self calibrating $\Rightarrow$ better systematics control
- Q_{EC} and atomic de-excitation spectrum poorly known
- complex pile-up spectrum
- in case of higher $Q \Rightarrow$ less sensitive

LTD projects with ^{163}Ho :

- ECHo, MMC detectors (Heidelberg)
- HOLMES, TES detectors (Milano, Genova, LNGS, NIST)
- Los Alamos Nat. Lab., ...

Common technical challenges:

- Clean ^{163}Ho production
- ^{163}Ho incorporation
- Large channel number $\Rightarrow$ high speed MUX
- Data handling (processing, storage, ...)





First detector prototype

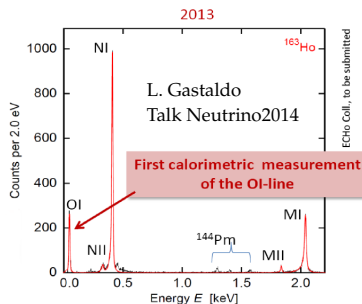
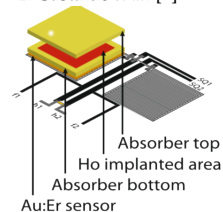
- low temperature metallic magnetic calorimeters
- embedding of ^{163}Ho source
 $\Rightarrow$ ion implantation @ ISOLDE-CERN
- about 0.01 Bq per pixel
- two pixels have been simultaneously measured

Calorimetric spectrum

- rise time ≈ 130 ns
- $\Delta E_{\text{FWHM}} = 7.6\text{eV}$ @ 6 keV
- non-linearity $< 1\%$ @ 6 KeV
- presently most precise ^{163}Ho spectrum

	$E_{\text{H lit.}}$	$E_{\text{H exp}}$	$\Gamma_{\text{H lit}}$	$\Gamma_{\text{H exp}}$
MI	2.047	2.040	13.2	13.7
MII	1.845	1.836	6.0	7.2
NI	0.420	0.411	5.4	5.3
NII	0.340	0.333	5.3	8.0
OI	0.050	0.048	5.0	4.3

L. Gastaldo *et al.* [7]



P.C.O Ranitzsch *et al* [8] L. Gastaldo *et al* [9]



^{163}Ho source - (n, γ) - reaction on ^{162}Er

- June 2012 : one irradiation at BER II Research Reactor Berlin
- Summer 2013: Two irradiations at ILL
 - treatment of Er prior to irradiation: all elements lighter than Er separated
 - treatment of Er after irradiation: all elements heavier than Ho are separated

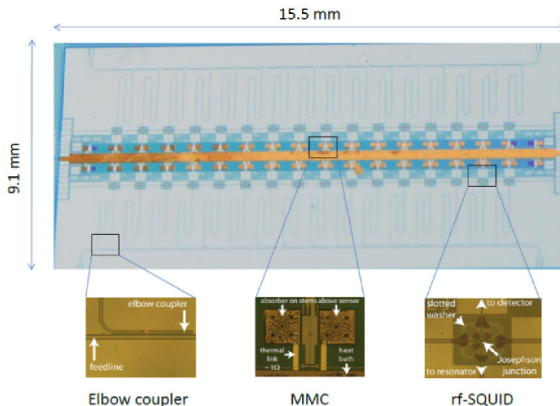
^{163}Ho source purification

- Radionuclides contained in the samples: ^{163}Ho and $^{166\text{m}}\text{Ho}$
- Mass separation/ion implantation
 - Separation at CERN/ISOLDE: Proposal accepted, will run in 2014
 - Use of the RISIKO mass-separator at Uni-Mainz, Summer 2014



Next step

- Microwave multiplexing technique
- 2 chips $\rightarrow$ 64 pixels
- $\Delta E_{FWHM} = 5\text{ eV}$, 10 Bq/pixel
- statistical sensitivity as low as 10 eV
- prove scalability





Goal

- Neutrino mass measurement:
⇒ m_ν statistical sensitivity as low as 0.4 eV
- Demonstrate technique potential and its scalability (Megapixel experiment)
- assess EC Q-value
- assess systematic errors

Baseline

- Transition Edge Sensors (TES) with ^{163}Ho implanted Bi:Au absorbers
- $6.5 \cdot 10^{13}$ nuclei per detector ⇒ 300 dec/s
- $\Delta E \simeq 1$ eV and $\tau_{\text{rise}} \simeq 1$ μs
- 16 channel demonstrator data taking starting from the second half of 2016
- 1000 channel final array from second half of 2017
 - $6.5 \cdot 10^{16}$ ^{163}Ho nuclei ⇒ 18 μg
 - $3 \cdot 10^{13}$ events in 3 years

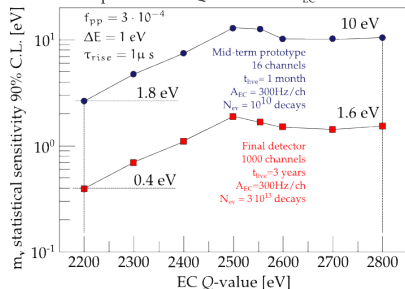


HOLMES tasks

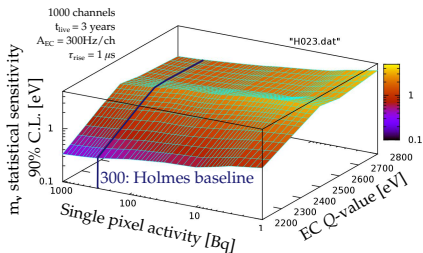
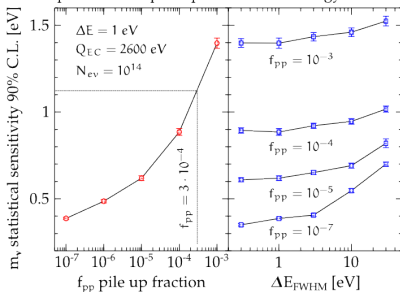
- ^{163}Ho isotope production and ^{163}Ho isotope embedding in detector
- Single TES optimization and testing
- TES array design, engineering and testing
- SQUID read-out and multiplexing optimization and testing (rf-SQUID)
- Real time and offline signal processing and analysis (trigger, OF filter, pile-up rejection)



^{163}Ho decay experiments statistical sensitivity dependence on Q-value with $A_{\text{EC}} = 300 \text{ Hz}$



^{163}Ho decay experiments statistical sensitivity dependence on pile-up fraction and energy resolution

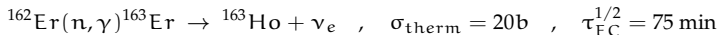


- The experimental sensitivity is directly related to $f_{\text{pp}} = \tau_{\text{rise}} \cdot A_{\text{EC}}$ (pile-up)
- The impact of the energy resolution is relatively smaller than that of pile-up
- In presence of a high level of pile-up, the experiment is relatively less sensitive to the energy resolution
- A constant background b is negligible if it is much smaller than the pile-up spectrum

^{163}Ho production and embedding

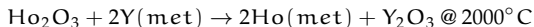


Production of ^{163}Ho : neutron activation of enriched ^{162}Er



- requires ^{162}Er enrichment and oxide chemical form (Er_2O_3)
- high yield (not all cross sections are measured)
- unavoidable $^{166\text{m}}\text{Ho}$: $^{165}\text{Ho}(n, \gamma)^{166\text{m}}\text{Ho}$ ($\beta_{12} = 1200\text{ y}$) from Ho contaminations or $^{164}\text{Er}(n, \gamma)^{165}\text{Er}$
- unavoidable ^{164}Ho : fast neutron activation $^{163}\text{Ho}(n, \gamma)^{164}\text{Ho}$
- nuclear reactor at the Institut Laue Langevin (ILL) in Grenoble

Thermoreduction to obtain the metallic Ho target for implantation:

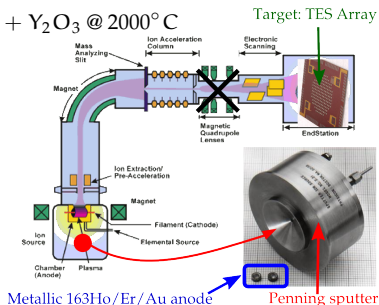


^{163}Ho separation from Dy, Er, Tm and Tr

- radiochemistry (before and/or after irradiation) in collaboration with PSI
- magnetic mass separation

^{163}Ho embedding in detector absorber

- implantation + magnetic separation
- Au film seposition for full containment





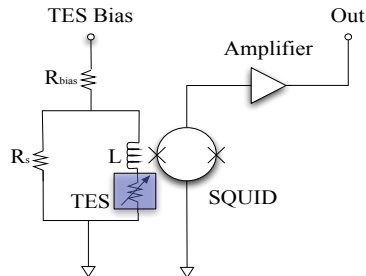
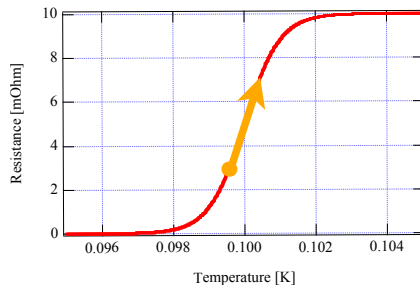
Transition Edge Sensor (TES): cryogenic particle detector that exploits the strongly temperature-dependent resistance of the superconducting phase transition.

Why TES?

- Widely used by the X-ray astrophysics community over the past twenty years
- Excellent energy resolution:

$$\Delta E_{FWHM} = \begin{cases} 1.26 \text{ eV @ } 1.5 \text{ keV} \\ 1.58 \text{ eV @ } 6 \text{ keV} \\ 1.94 \text{ eV @ } 8 \text{ keV} \end{cases}$$

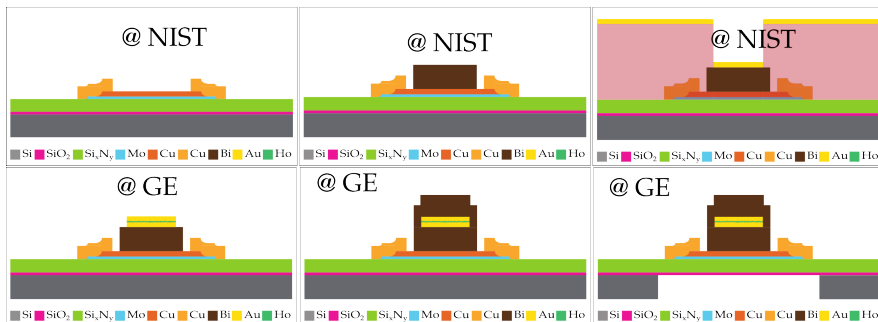
- Low impedance
- Large Array and multiplexing (TDM, CDM, FDM, μmux)
- Tunable critical temperature T_C exploiting the proximity effect [10] $\Rightarrow$ Mo: Au or Mo: Cu proximity TES ($T_C \simeq 100 \text{ mK}$)





Transition Edge Sensor (TES): MoCu bilayers $\rightarrow T_c \approx 100\text{mK}$

- microcalorimeters with electro-thermal feedback
- $3\text{ }\mu\text{m}$ thick Bi absorber with ^{163}Ho /Au encapsulating for full absorption
- source: thin electrodeposited Au encapsulating implanted ^{163}Ho
- TES fabricated by subcontractor (NIST, Boulder, CO, USA)
- ^{163}Ho implantation and Si_2N_3 membrane release at Genova



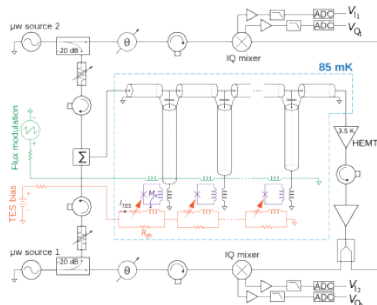
rf-SQUID read-out and ROACH2



Microwave SQUID multiplexing (μ Mux) is a read-out technique that combines the proven sensitivity of TESs and the scalable multiplexing power found in KIDs. It uses radio-frequency rf-SQUIDs coupled to high quality-factor microwave resonators.

rf-SQUID read-out

- DC biased TES
- SQUID coupled with TES and a resonator circuit
- Microwave rf-SQUID read out with flux ramp modulation (common flux line is inductively coupled to all the SQUIDs)
- Signal reconstructed by homodyne detection and demodulation



ROACH2-based Software Defined Radio

- Xilinx FPGA based digital data processing
- frequency comb generation (up to ≈ 60 in 0-550 MHz)
- GHz band up/down conversion (5-5.5 GHz)
- homodyne detection: IQ signal de-multiplexing
- signal channelizing and rf-SQUID signal de-modulation
- real time signal processing $\rightarrow$ 140TB in 3 years

IF Board

ADC/DAC Board

ROACH2



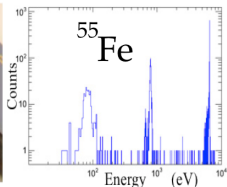
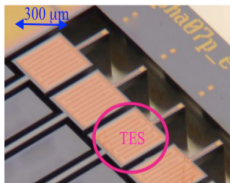
^{163}Ho source

- Proton irradiation of natural Dy at high-current LANL proton accelerator
- Separation of ≈ 500 μg di Ho from 110 mg di Dy

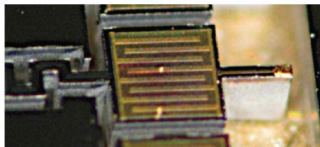
J.W. Engle *et al.* [11]

Detector

- Transition Edge Sensor: MoCu superconducting films on solid silicon



- Feasibility demonstrated with embedding ^{55}Fe source
- Recent measurements show $\Delta E \approx 7.5$ eV FWHM
- Testing several methods of incorporating Ho into absorbers



from L. Gastaldo - talk at Neutrino2014 Conference



- The measurement of the end point of nuclear beta or electron capture (EC) decays spectra is the only model-independent
- The goal of the next generation experiments is the sub-eV neutrino mass sensitivity
- ECHo has acquired the most precise ^{163}Ho spectrum so far
- The ECHo next step is to achieve a sensitivity on neutrino mass of the order of 10 eV
- The HOLMES experiment will perform a direct measurement of the neutrino mass by using microcalorimeter with ^{163}Ho -implanted absorber.
- The Goals of the HOLMES experiment are:
 - assess EC Q-value of the ^{163}Ho
 - assess systematic errors
 - achieve a statistical sensitivity as low as 0.4 eV
 - prove technique potential and its scalability (Megapixel experiment)



- [1] **A. De Rujula and M. Lusignoli** *Physics Letters B*, vol. 118, pp. 429–434, 1982, doi:10.1016/0370-2693(82)90218-0.
- [2] **A. de Rujula and M. Lusignoli** *Nuclear Physics B*, vol. 219, pp. 277 – 301, 1983, doi:10.1016/0550-3213(83)90642-9.
- [3] **C.W. Reich and Balraj Singh** *Nuclear Data Sheets*, vol. 111, pp. 1211 – 1469, 2010, doi:10.1016/j.nds.2010.04.001.
- [4] **G. Audi, A.H. Wapstra and C. Thibault** *Nuclear Physics A*, vol. 729, pp. 337 – 676, 2003, doi:j.nuclphysa.2003.11.003.
The 2003 NUBASE and Atomic Mass Evaluations.
- [5] **A. Nucciotti** 2014, e-print: [arXiv:1405.5060](#) [physics.ins-det].
- [6] **A. Nucciotti, E. Ferri and O. Cremonesi** *Astropart.Phys.*, vol. 34, pp. 80–89, 2010, doi:10.1016/j.astropartphys.2010.05.004, e-print: [arXiv:0912.4638](#) [hep-ph].
- [7] **L. Gastaldo et al.** *Journal of Low Temperature Physics*, 2014, doi:10.1007/s10909-014-1187-4.
- [8] **P.C.O. Ranitzsch et al.** *Journal of Low Temperature Physics*, vol. 167, p. 10004, 2012.
- [9] **L. Gastaldo et al.** *Nuclear Inst. Meth.*, vol. 711, pp. 150–159, 2013.



- [10] **Hans Meissner** *Physical Review*, vol. 117, pp. 672–680, February 1960,
[doi:10.1103/PhysRev.117.672](https://doi.org/10.1103/PhysRev.117.672).
- [11] **J.W Engle et al.** *NIM B*, vol. 311, pp. 131–138, 2013.