

# A direct neutrino mass measurement with $^{163}\text{Ho}$ : the **HOLMES** experiment

M. De Gerone

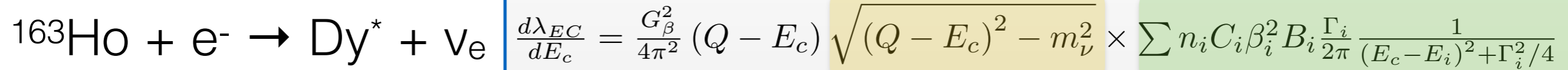
INFN Genova

on behalf of the HOLMES collaboration

XX Lomonosov Conference on elementary particle physics

Moscow State University, August 19-25 2021

# $^{163}\text{Ho}$ electron capture



**Q~2.8keV**, capture only from shell  $\geq$  M1

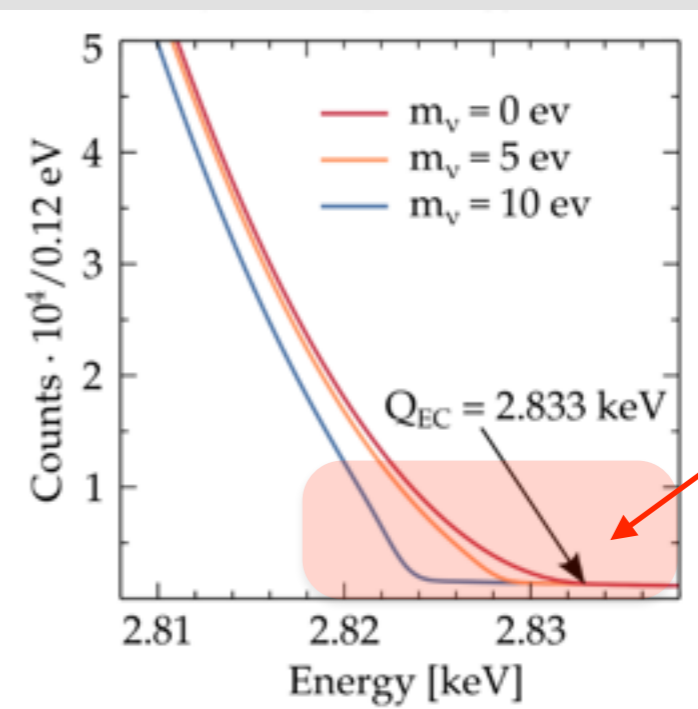
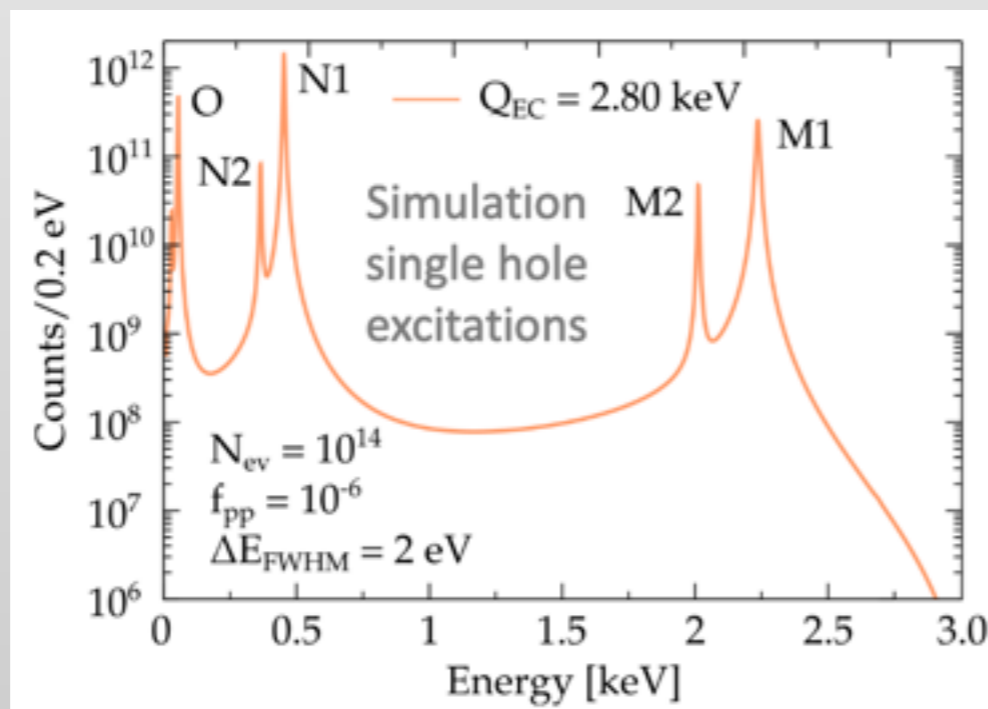
De Rujula & Lusignoli, Phys. Lett. B 118 (1982) 429

**same factor as  $\beta$  decay**

(total de-excitation energy  $E_c$  instead of  $E_e$ )

Breit-Wigner shape

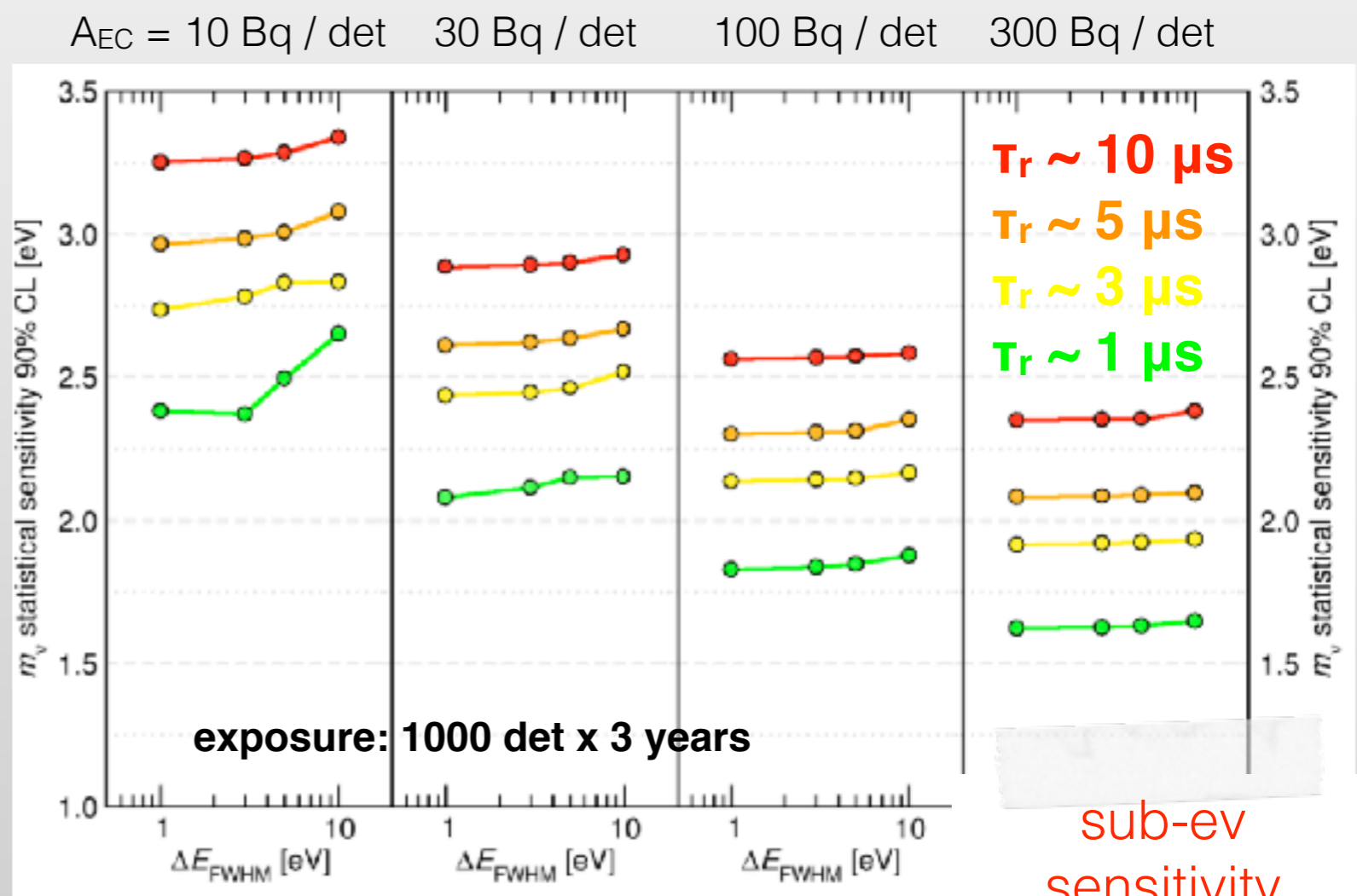
- calorimetric measurement of  $\text{Dy}^*$  de-excitation
- “good” event rate and  $\nu$  mass sensitivity depends on Q-value and capture peak position (roughly  $\sim 1/(Q - E_{M1})^3$ )
- $\tau_{1/2} \sim 4570$  years  $\rightarrow$  few active nuclei needed



expected  $\nu$  mass effect

# The **H<sub>V</sub>LMES** project in a nutshell

- **Direct neutrino mass measurement** with statistical sensitivity around 1 eV
- Usage of **Transition Edge Sensor (TES) based micro-calorimeters** with  $^{163}\text{Ho}$  implanted Au absorber:
  - $6.5 \times 10^{13}$  nuclei / det
  - $A_{\text{EC}} \sim 300$  Bq / det
  - $\Delta E \sim 1$  eV,  $\tau \sim 1\text{-}10$   $\mu\text{s}$
- **2 steps approach:**
  - 64 channels prototype,
    - $t_m = 1$  month,  $m_\nu < 10$  eV
  - 1000 channels array:
    - $6.5 \times 10^{16}$  total nuclei
    - $O(10^{13})$  events / year
    - $m_\nu \sim 1$  eV
- Should prove the technique potential and scalability by evaluating EC spectral shape and systematic errors



sub-ev  
sensitivity  
region:  
 $N_{\text{ev}} \gg 10^{13}$

# $^{163}\text{Ho}$ electron capture

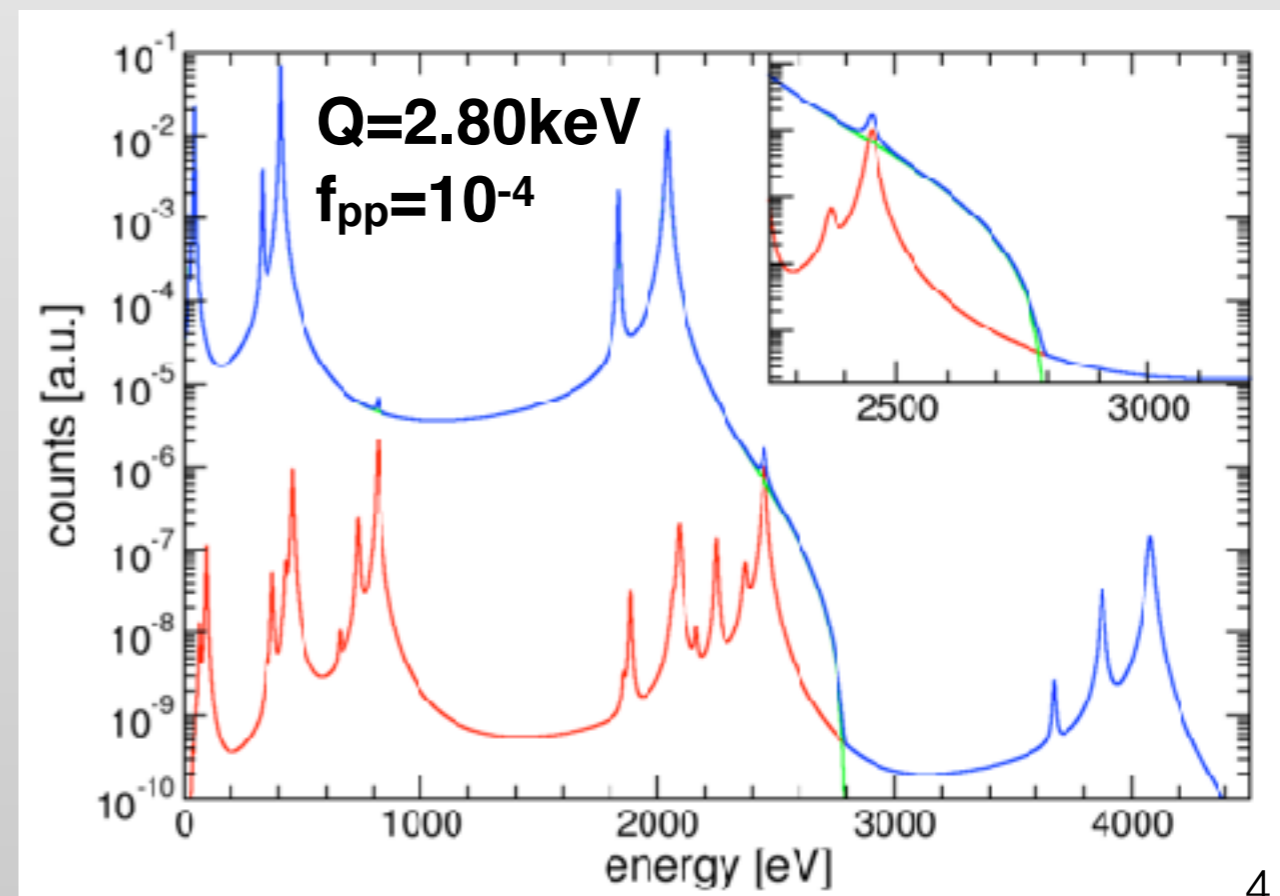
**Complex pile-up spectrum:** end-point is dominated by  $((Q-E_c)^2 - m_\nu^2)^{1/2}$  but expected distortions due to pile-up:

$$N_{pp}(E) = f_{pp} N_{EC}(E) \otimes N_{EC}(E)$$

Pile-up occurs when multiple events arrive within the resolving time of the detector. In a first approximation, fraction of unresolved pile up is given by  $f_{pp} = \tau \times A_{EC}$ .

In order to reduce pile-up:

- trade-off between activity and statistic;
- detector with fast signal rise time  $\tau_r$ ;
- pile-up resolving algorithm.



# $^{163}\text{Ho}$ production

$^{163}\text{Ho}$  does not exist in nature: it is produced from  $^{162}\text{Er}$  neutron activation at nuclear reactor:

- $^{162}\text{Er} (n,\gamma) ^{163}\text{Er}$ ,  $\sigma_{\text{therm}} \sim 20 \text{ b}$
- $^{163}\text{Er} + e^- \rightarrow ^{163}\text{Ho} + \nu_e$  ( $\tau_{1/2} \sim 75 \text{ m}$ )
- high yield
  - $\sim 3 \times 10^{12}$   $^{163}\text{Ho}$  nuclei/mg  $^{162}\text{Er}$ /h
- requires  $^{162}\text{Er}$  enrichment and oxide chemical form ( $\text{Er}_2\text{O}_3$ )

It is a “dirty” process: many other isotopic species are created together with  $^{163}\text{Ho}$ . Worst one:

- $^{165}\text{Ho} (n,\gamma) ^{166\text{m}}\text{Ho}$  ( $\beta$ ,  $\tau_{1/2} \sim 1200 \text{ years}$ )
- from Ho contamination or  $^{164}\text{Er} (n,\gamma)$
- need high purification of sample:
  - radiochemical separation:
    - removes everything but Ho;
  - mass separation with magnetic dipole.

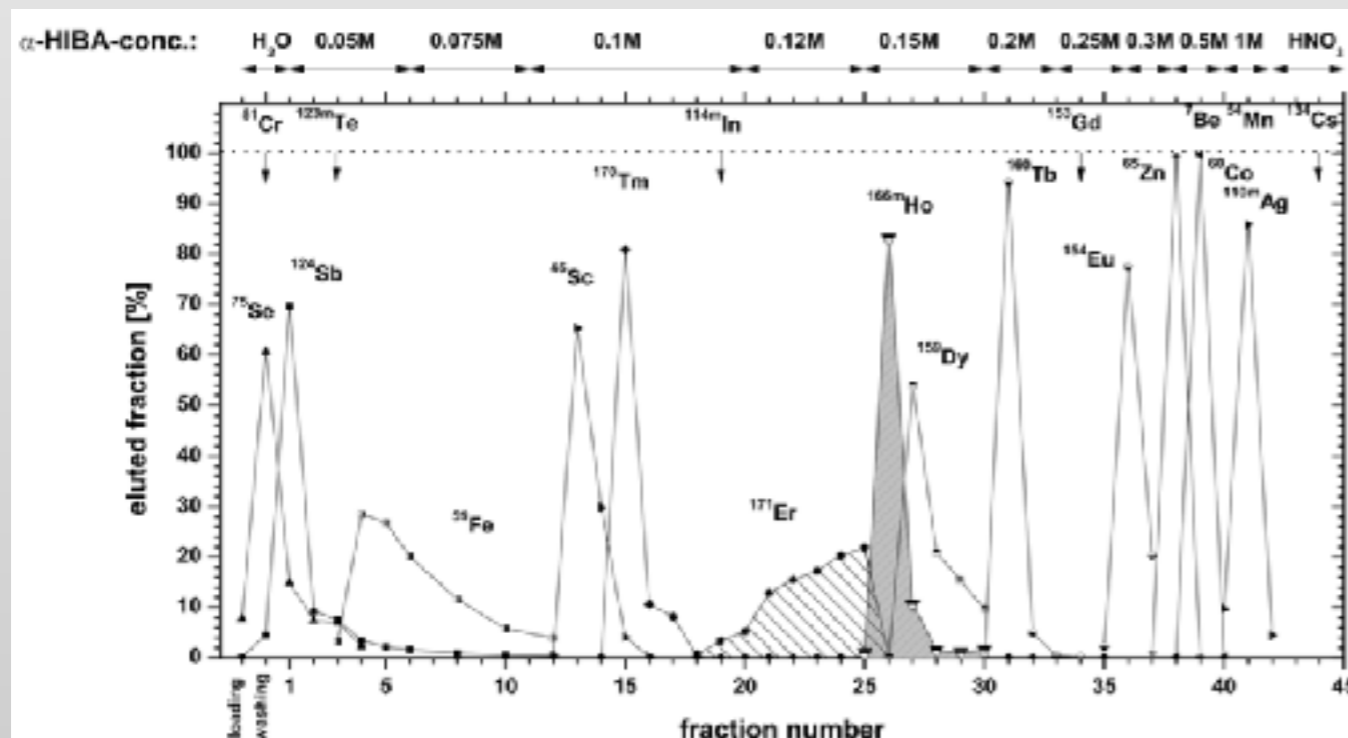
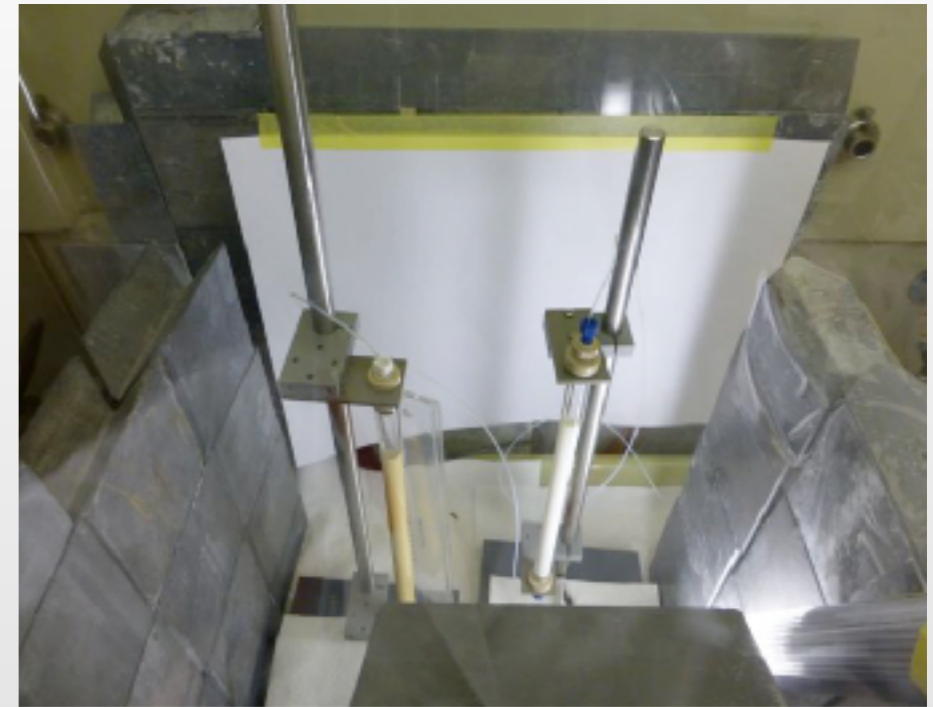
Tm 163 1.81 h	Tm 164 5.1 m / 2.0 m	Tm 165 30.06 h	Tm 166 7.70 h	Tm 167 9.25 d	Tm 168 93.1 d
Er 162 0.139	Er 163 75 m	Er 164 1.601	Er 165 10.3 h	Er 166 33.503	Er 167 2.3 s / 22.869
Ho 161 6.7 s / 2.5 h	Ho 162 68 m / 15 m	Ho 163 1.1 s / 4570 a	Ho 164 37 m / 29 m	Ho 165 100	Ho 166 1200 a / 28.80 h
Dy 160 2.329	Dy 161 18.889	Dy 162 25.475	Dy 163 24.896	Dy 164 28.260	Dy 165 1.3 m / 2.35 h
Tb 159	Tb 160	Tb 161	Tb 162	Tb 163	Tb 164



In collaboration with  
ILL (Grenoble, FR)  
Paul Scherrer Institute (Villigen, CH)

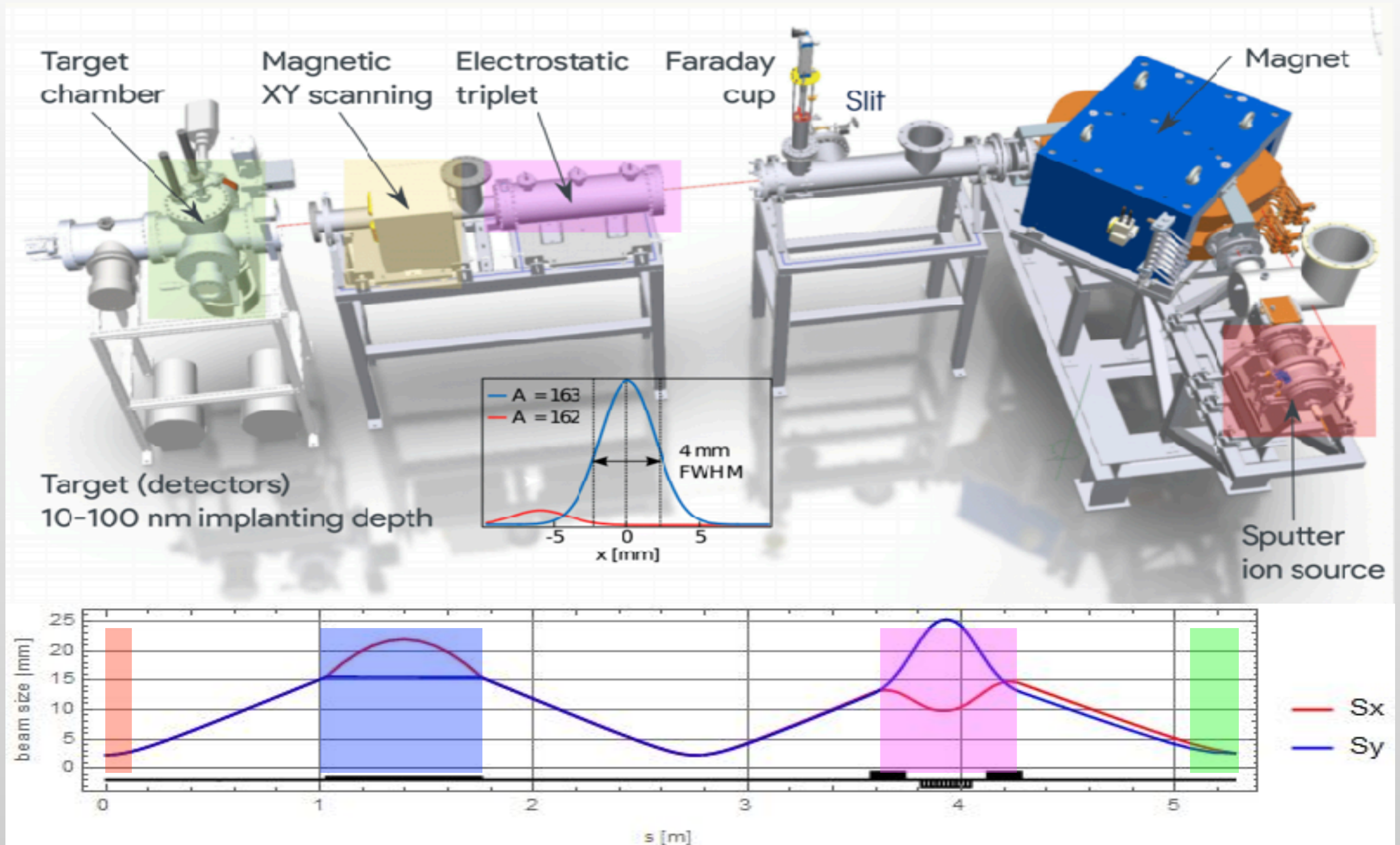
# $^{163}\text{Ho}$ purification

- Enriched  $\text{Er}_2\text{O}_3$  samples irradiated at ILL (Grenoble) and post-processed at PSI:
  - 25 mg, 55 days irradiation  $\rightarrow A(^{163}\text{Ho}) \sim 5 \text{ MBq}$
  - 150 mg, 50 days irradiation  $\rightarrow A(^{163}\text{Ho}) \sim 38 \text{ MBq}$
  - 540 mg, 50 days irradiation  $\rightarrow A(^{163}\text{Ho}) \sim 120 \text{ MBq}$
- Ho radiochemical separation is performed via ion-exchange chromatography at PSI.
  - Expected  $^{166\text{m}}\text{Ho}$  contamination fraction:  $\sim 10^{-3}$
  - Currently available activity is already enough for HOLMES purposes.



S. Heinitz et al. PLOS ONE 13(8):e0200910

# $^{163}\text{Ho}$ mass separation



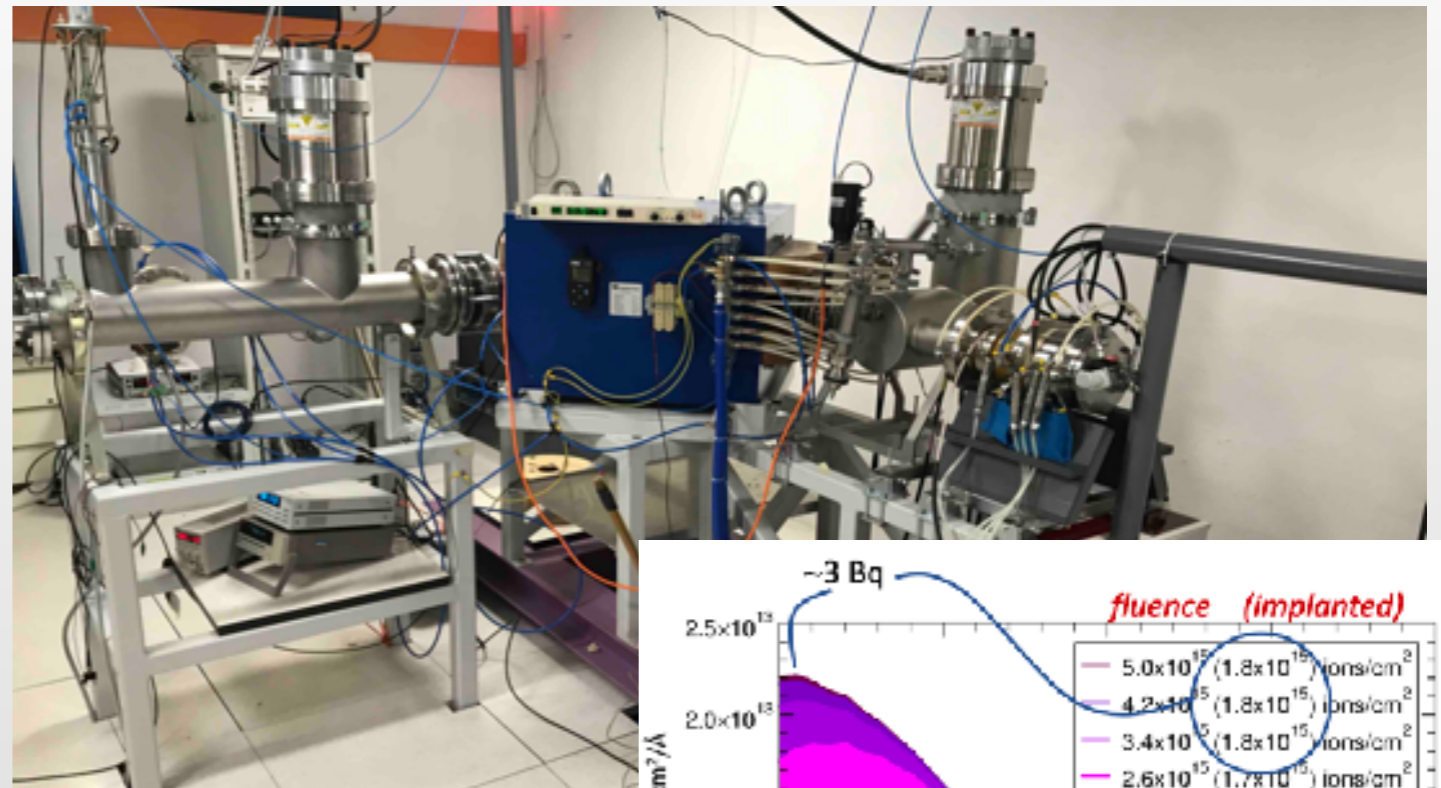
# $^{163}\text{Ho}$ mass separation

The layout currently installed in Genova's lab includes:

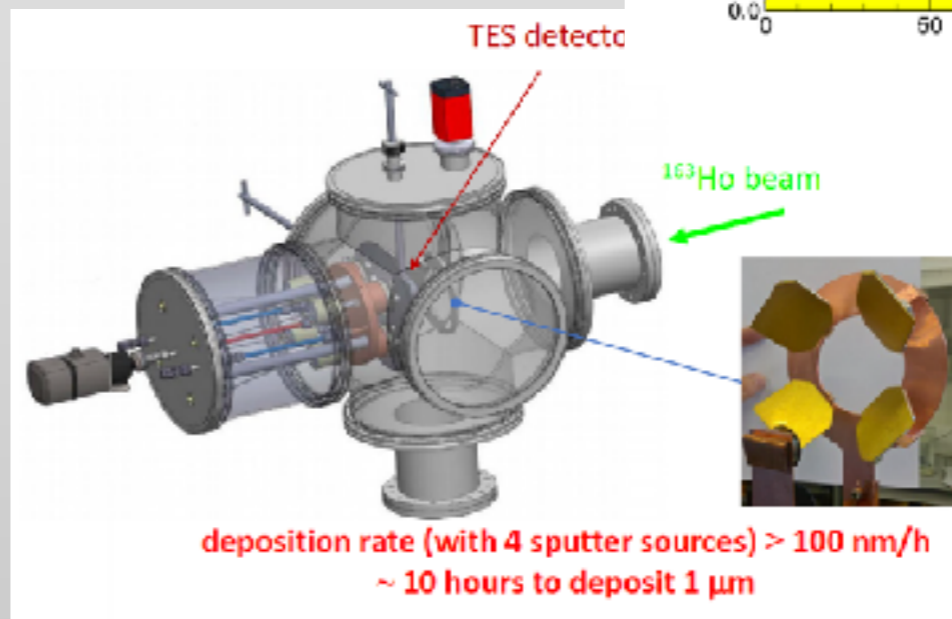
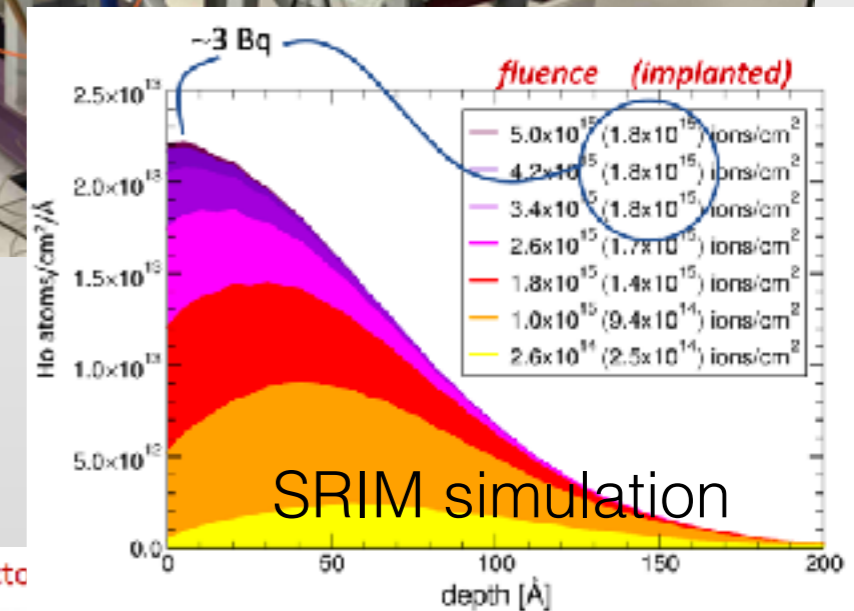
- an **argon Penning sputter ion source** with a 50kV acceleration section, corresponding to  $O(10\text{ nm})$  implantation depth;
- a **magnetic dipole mass analyzer**, field up to 1.1 T; expected  $^{163}/^{166}$  a.m.u. separation  $> 5\sigma$ ;
- a **Faraday cup and a slit**.

The **target chamber**, which allows simultaneous co-evaporation of Au will be installed soon. Co-evaporation is needed:

- to fully encapsulate the source in the absorber and
  - because after a while  $^{163}\text{Ho}$  concentration in the absorber saturates.
- The machine will be upgraded after first 64 detector arrays production.



Target chamber



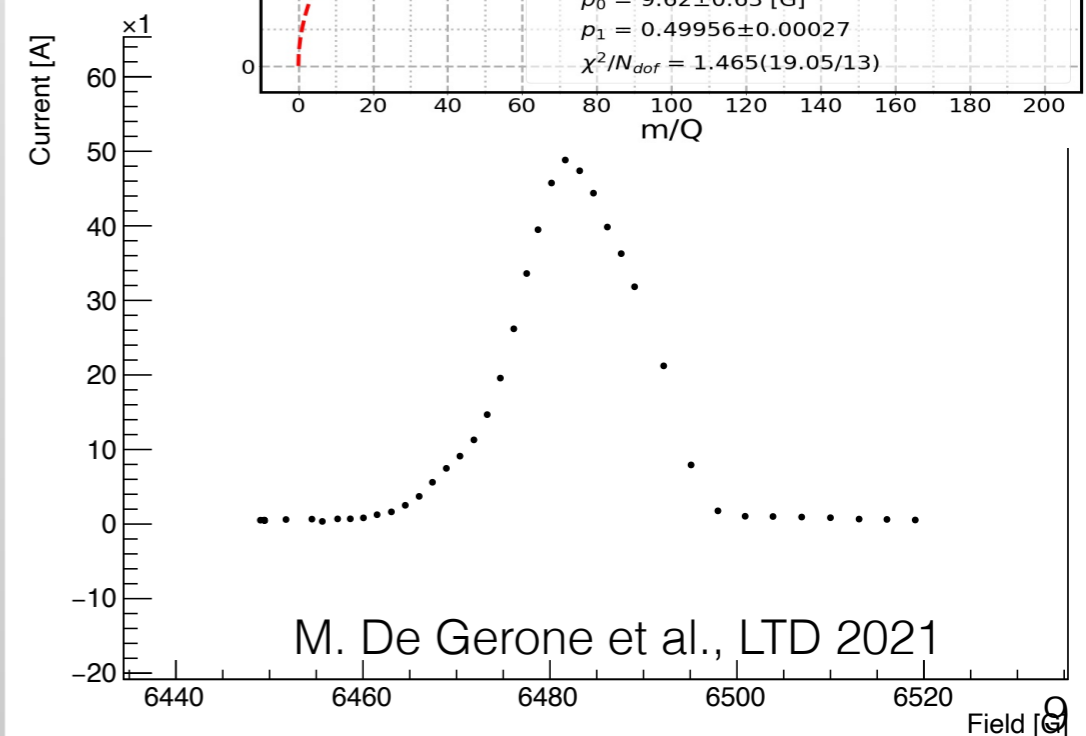
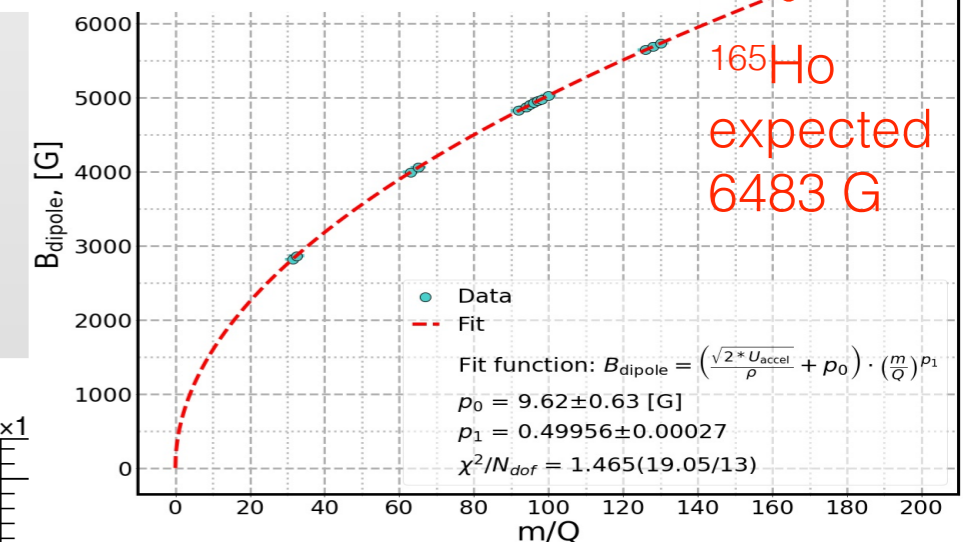
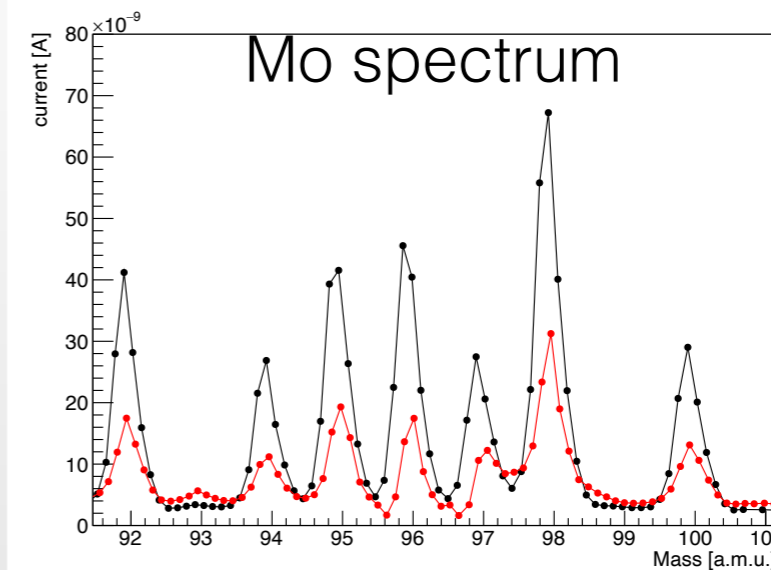


# Ion implanter calibration and test

The machine is calibrated using peaks from Cu, Au and Mo (from sputter target and anode). A small misalignment has been measured and taken into account. Mass resolving power has been evaluated from Cu and Mo peaks and extrapolate to be 18 mm at 163/166 a.m.u.

First tests performed with sputter target made with bulk Cu disk with a thin layer of  $^{165}\text{Ho}(\text{OH})_3$  deposited via molecular plating:

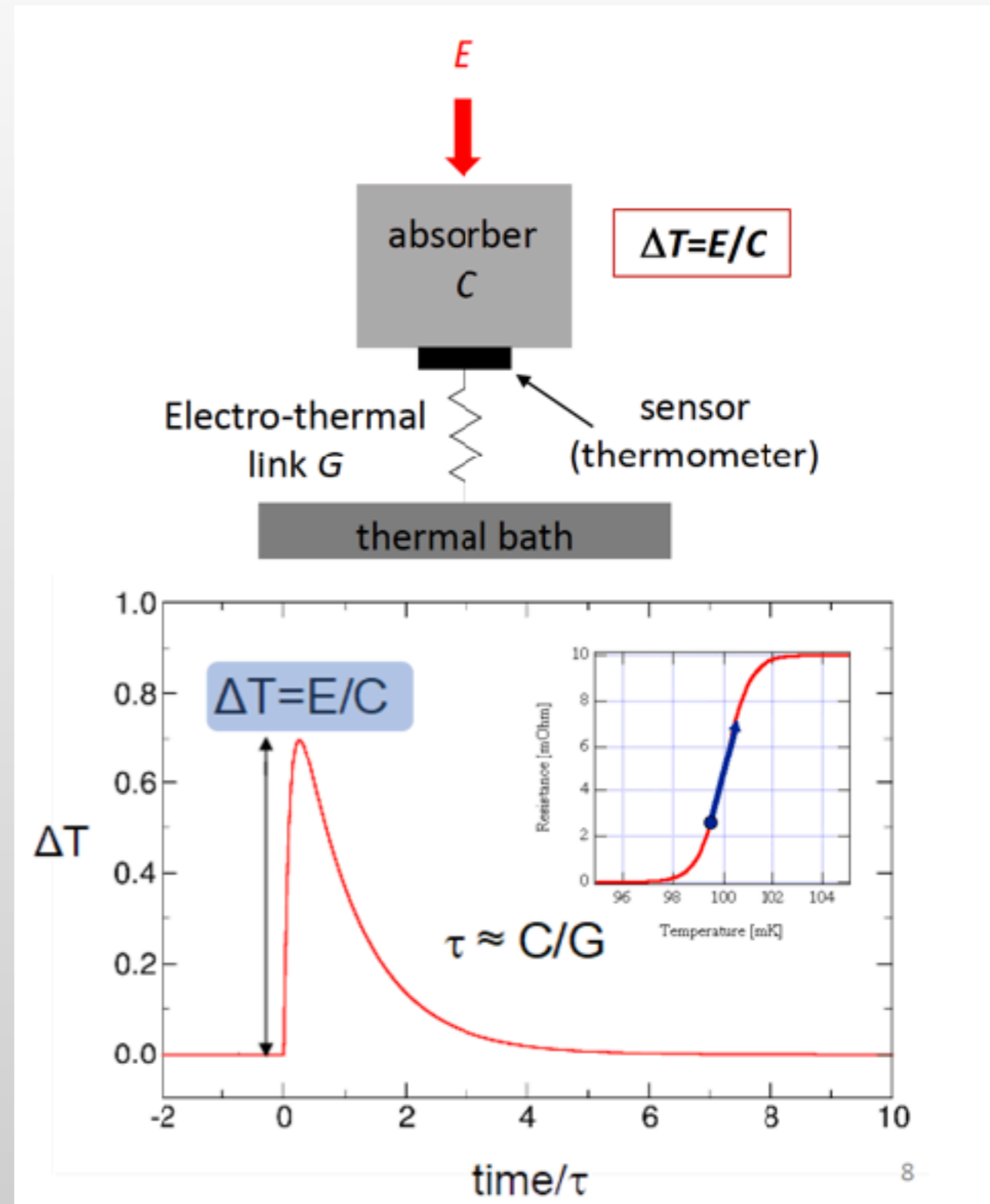
- clear peak at 165 a.m.u. BUT source efficiency is still quite low:
  - current O(50nA) sustained for few minutes...
  - Maybe  $^{165}\text{Ho}(\text{OH})_3$  dissociates in different compound? (HoO, Ho(OH)...) )
- Now testing:
  - Different Ho compounds to be included in sputter target:
    - Ho(CH<sub>3</sub>COOH), HoPO<sub>4</sub>, Ho(NO<sub>3</sub>)<sub>3</sub>
  - Different techniques for sputter target production:
    - Sintered target (mixture of Ho/Ti/Ni/Sn)



# TES based $\mu$ -calorimeters

## Transition edge sensors based $\mu$ -calorimeters:

- absorber coupled to a superconductive sensor (thermometer) kept in the transition region;
- energy release in the absorber produces a temperature increase in thermometer and then a change in TES resistance;
- **Exploit the steepness of  $R(T)$  of a superconductor kept in its transition to measure  $\Delta E$ :**
  - state of art **energy resolution ( $O(eV)$ )**;
  - **multiplexing readout scheme available**;
  - **limited dynamics**: design has to be optimized for a specific application.
  - $\Delta T_{\max} = E/C$ ,  $C$  = thermal capacity
  - $\Delta T(t) = E/C e^{-t/\tau}$ ,  $\tau = C/G$ ,  $G$  is the thermal conductance

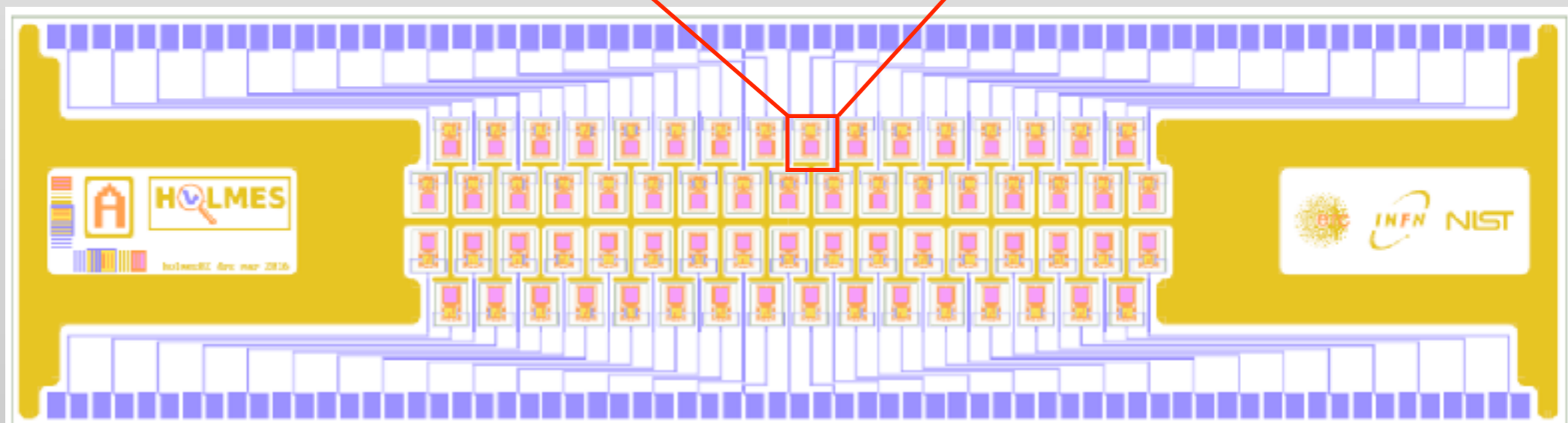
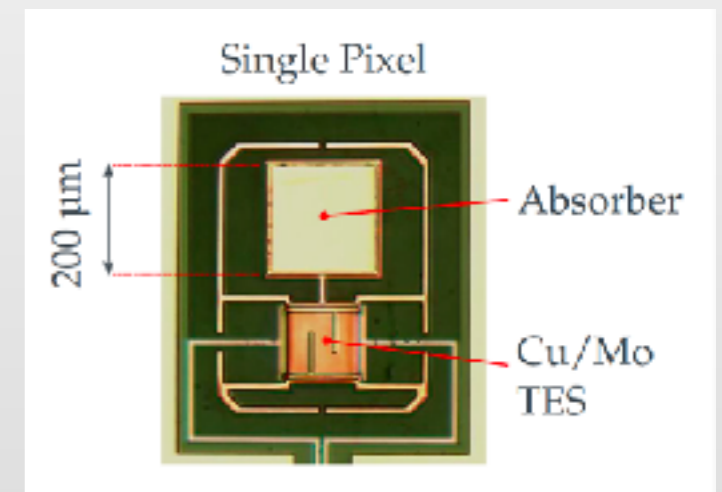
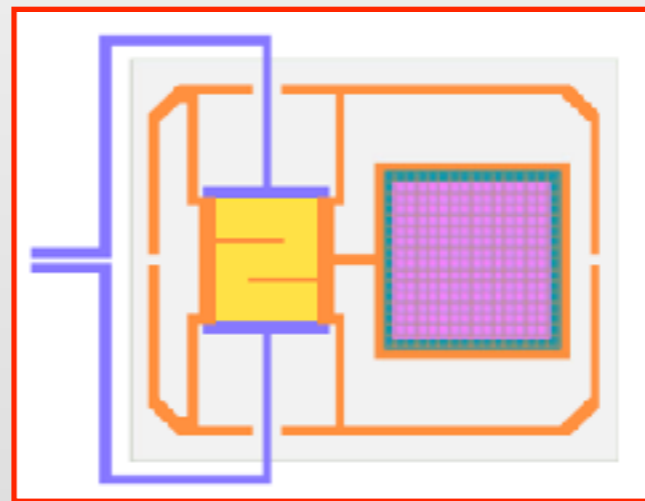


# Holmes detector design

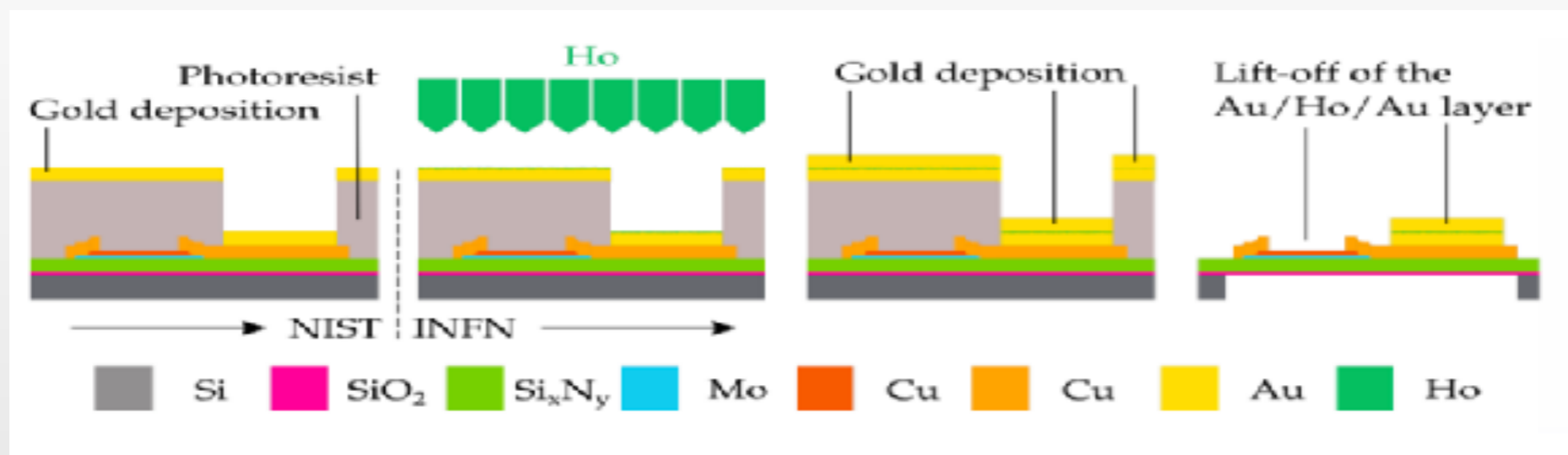
- TES design, production and preliminary test is done @NIST
- **2  $\mu\text{m}$  Au thickness for full absorption** of electrons and photons
- **“side car” configuration** to avoid TES proximization and allow G engineering for a better  $\tau$  control
- Design optimized to obtain best compromise between resolutions and time response.

Target performances (@ 3 keV):

- $\Delta E_{\text{FWHM}} \text{ O(eV)}$
- $\tau_{\text{rise}} \sim 10 \mu\text{s}$
- $\tau_{\text{decay}} \sim 100 \mu\text{s}$



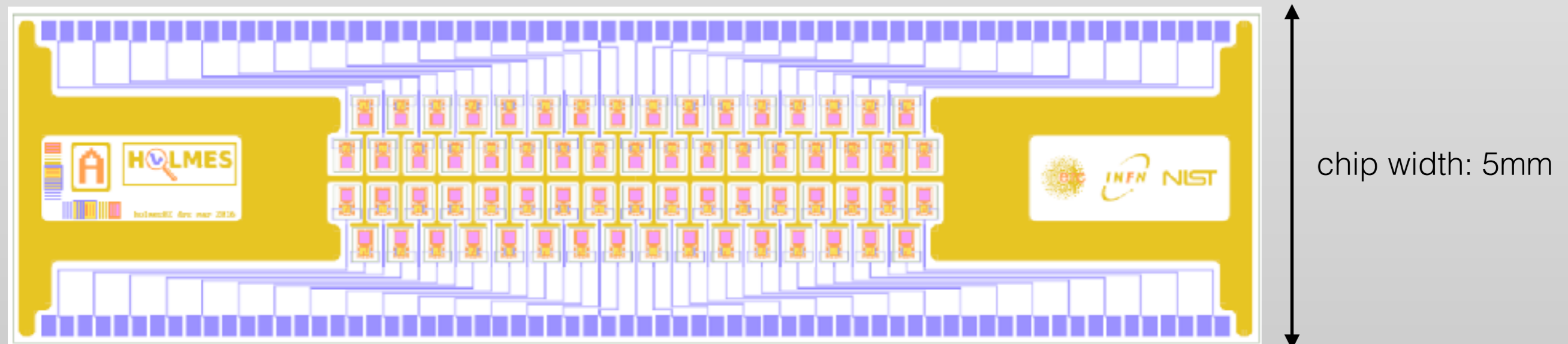
# Holmes Detector production



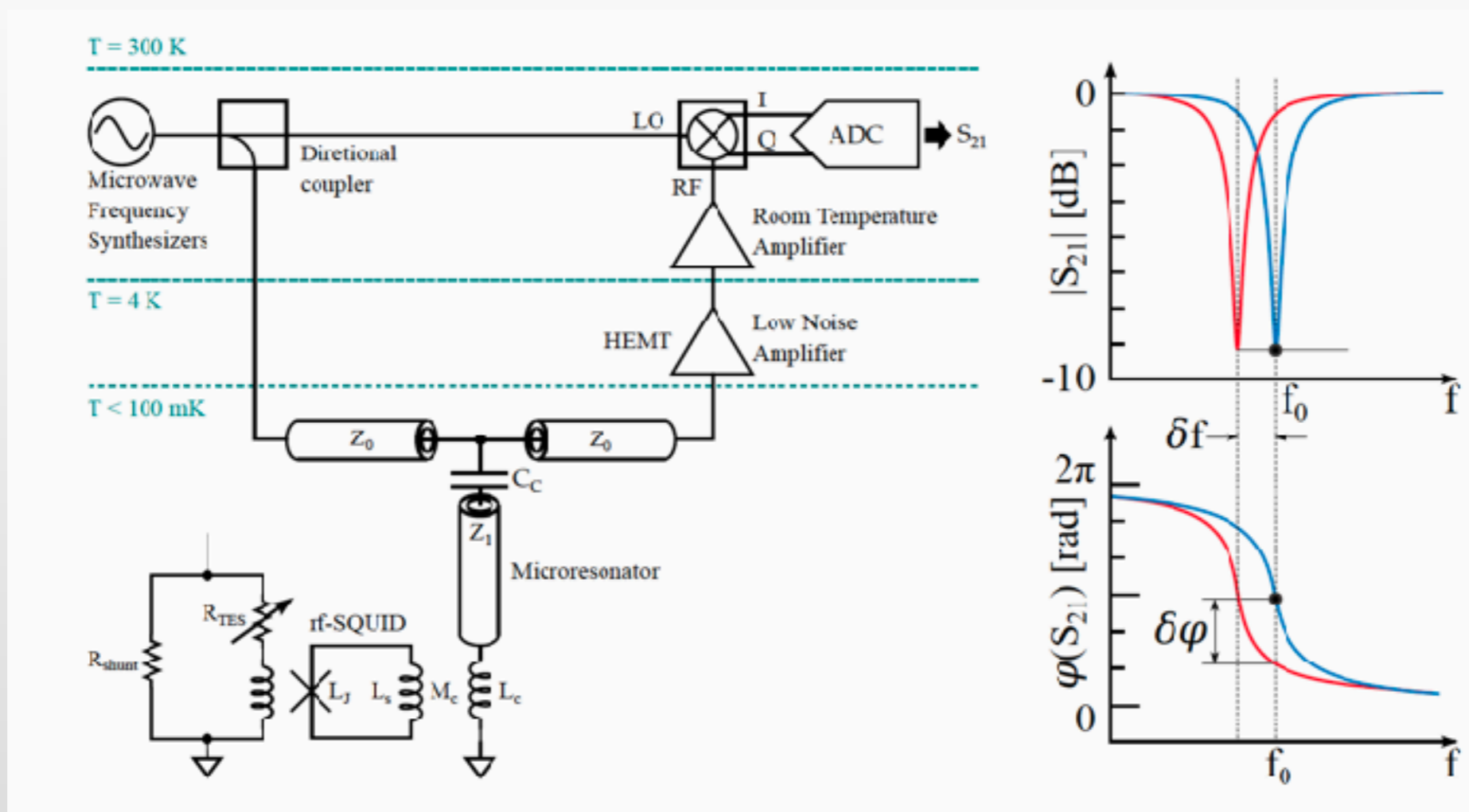
Detector fabrication is done with a **multi-step procedure**:

- 1) TES array is produced @NIST up to **first 1  $\mu\text{m}$  Au layer**;
- 2)  **$^{163}\text{Ho}$  implantation and Au co-evaporation**;
- 3) **1  $\mu\text{m}$  Au final layer** is deposited over Ho implantation (“complete” the absorber)
- 4) **membrane release** with KOH or DRIE process

4 x 16 linear array for implantation optimization and low parasitic L

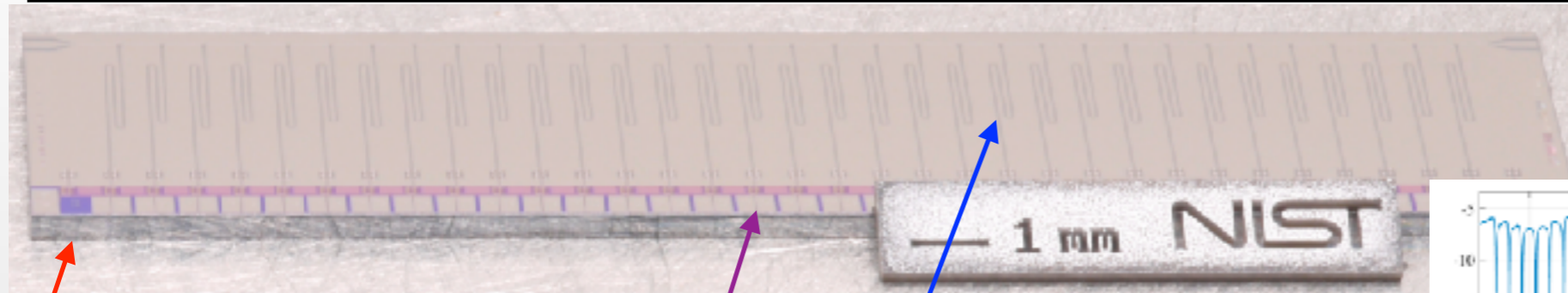


# Multiplexed readout

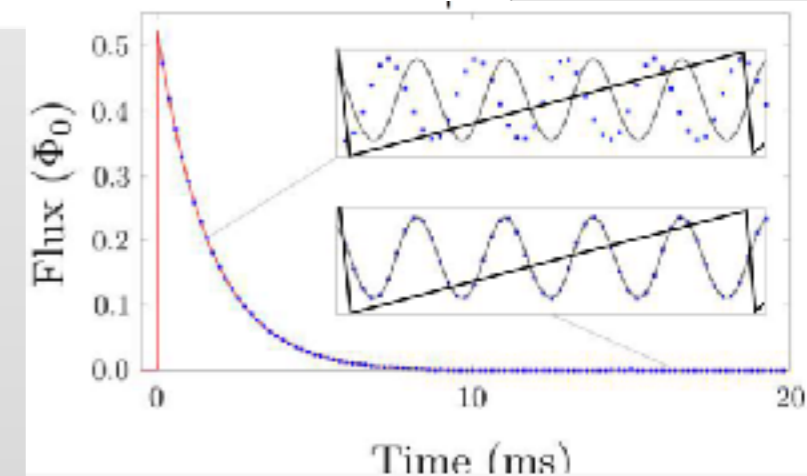
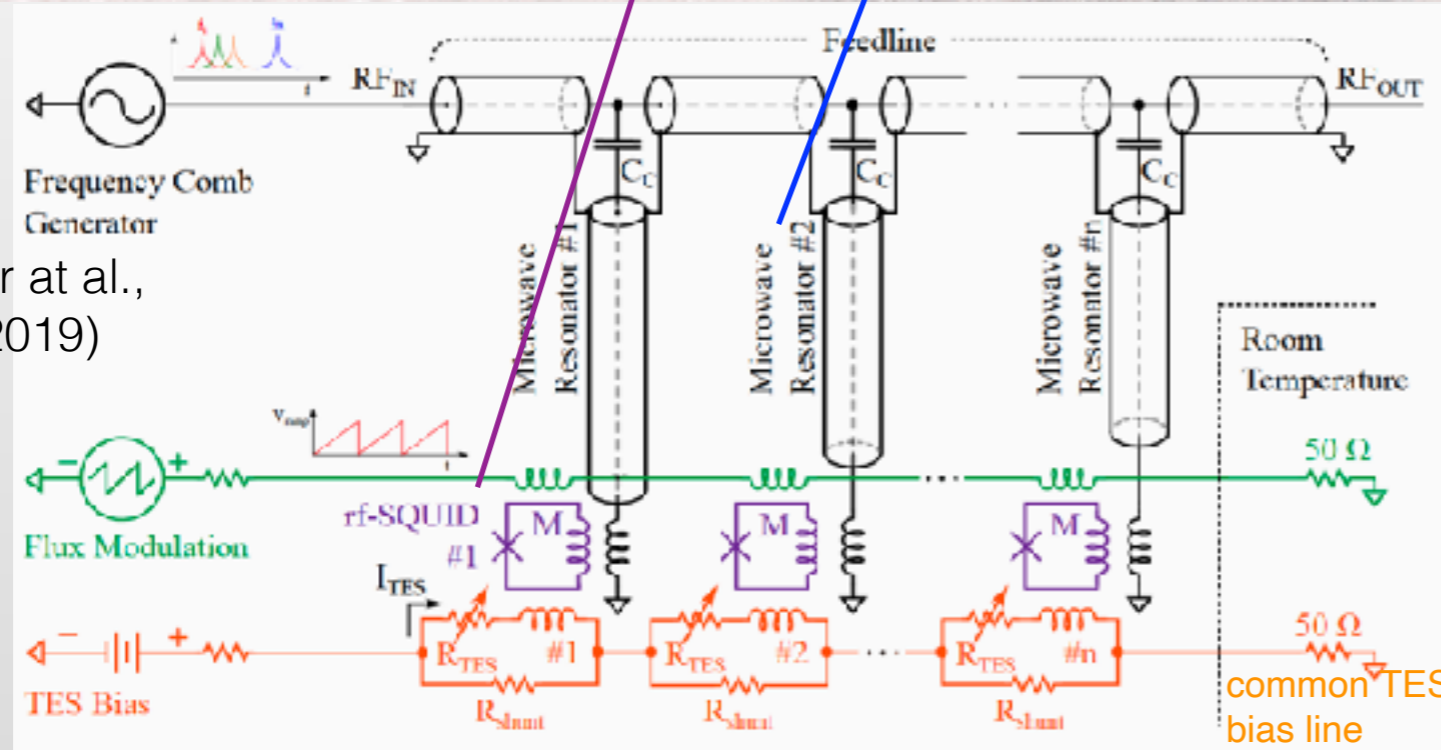
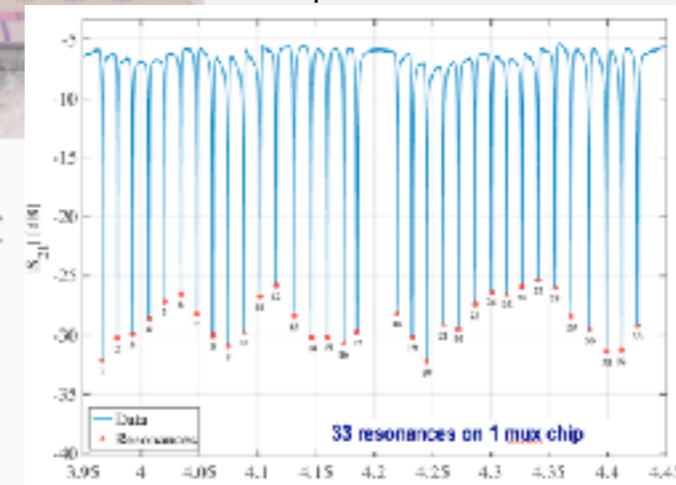


- DC-biased TES inductively coupled to a dissipation less RF-SQUID
- RF-SQUID inductively coupled to a high-Q superconducting  $\lambda/4$  resonator
- Change in TES current  $\Rightarrow$  change in the input flux to the SQUID
- Change in the flux to the SQUID  $\Rightarrow$  change of resonance frequency and phase
- Each micro-resonator can be continuously monitored by a probe tone

# Array readout: rf-SQUID $\mu$ wave mix



$\lambda/4$  resonators coupled to common feedline tuned at different frequencies for multiplexing

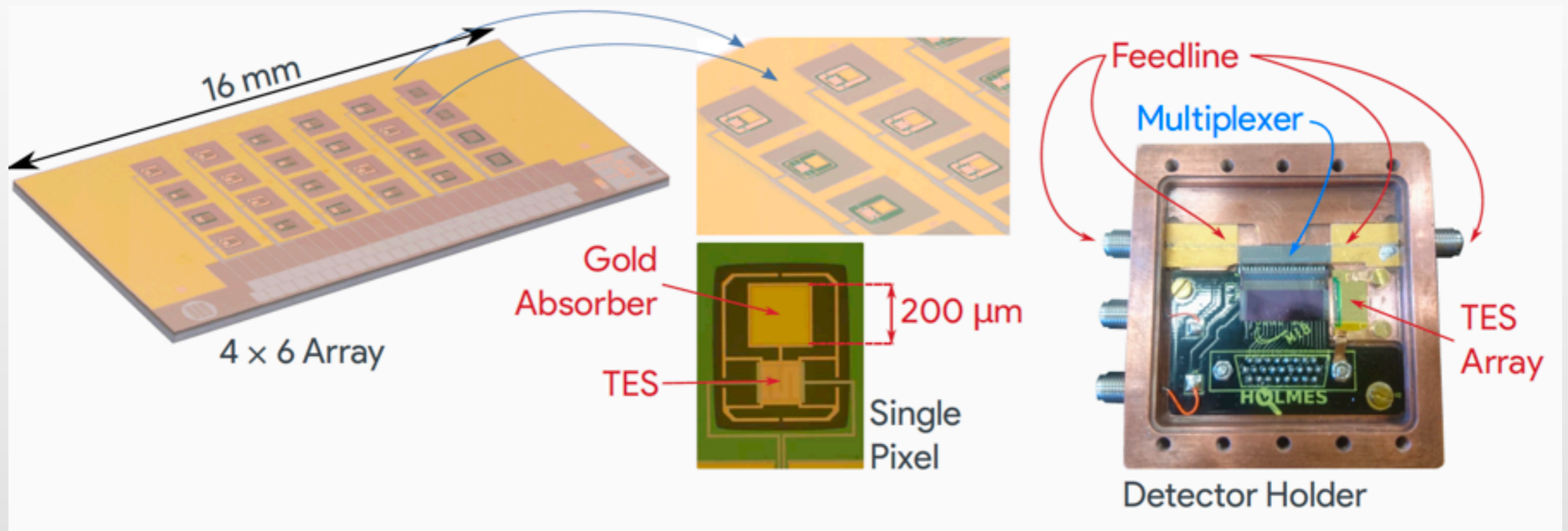


common SQUID modulation line

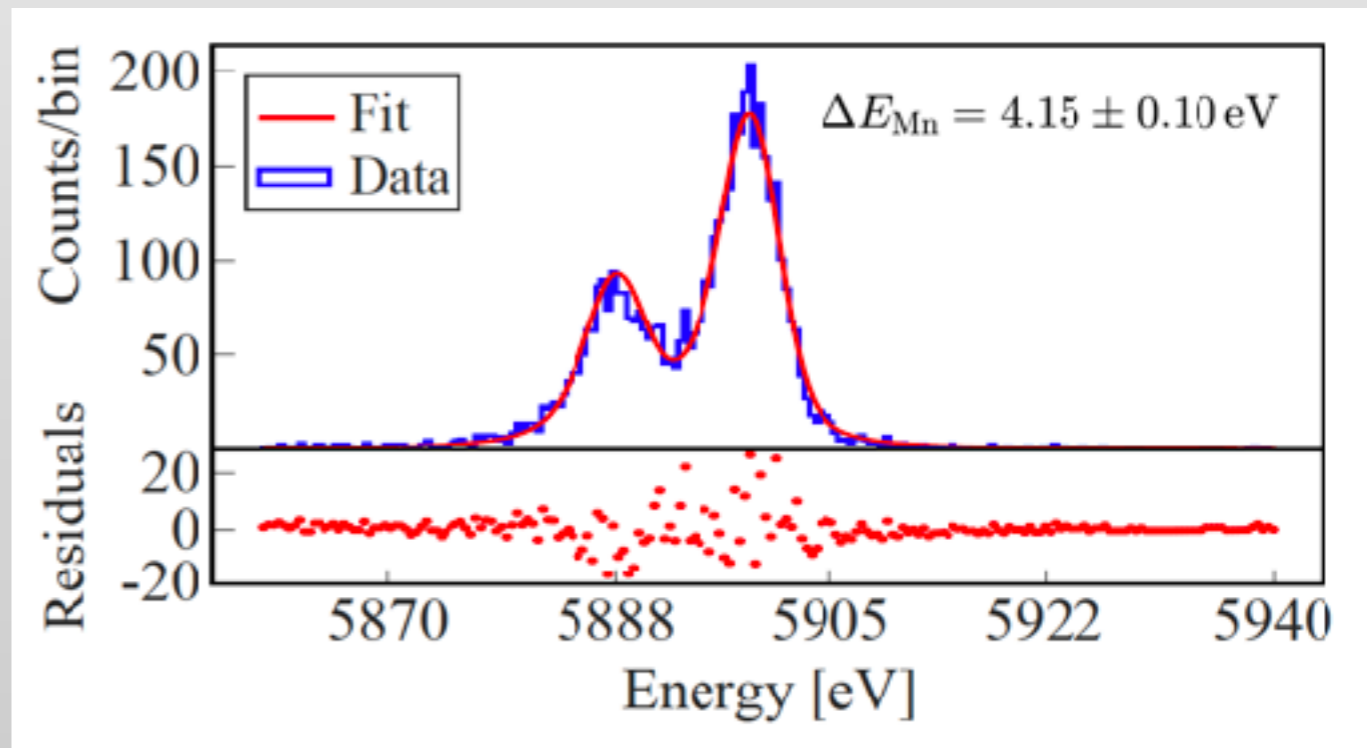
D. T. Becker et al.,  
JINST 14 (2019)  
10, P10035

- By coupling many resonators to a single microwave feedline it is possible to perform the readout of multiple detectors
- Sensors are monitored by a set of sinusoidal probe tones (frequency comb)
- The signal is reconstructed from the phase shift of the SQUID oscillation (solid line), with respect to a reference sine function (dotted line).
- The ramp induced a controlled flux variation in the rf-SQUID, crucial to linearize the response
- Large multiplexing factor ( $>100$ ), limited by digitizer bandwidth

# Pixel testing with HOLMES DAQ



- Tested different u-calorimeter geometries
- Produced at NIST
- Not yet implanted with Ho
- Sources:  $^{55}\text{Fe}$  + fluorescence sources (range 1.6 - 6 keV)
- Energy resolution  $\sim (4.5 \pm 0.3)$  eV
- Best performing detector:  $(4.15 \pm 0.10)$  eV @Mn  $K_{\alpha}$



# Conclusion and prospects

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- The HOLMES experiment aims to perform a direct neutrino mass measurement with calorimetric technique exploiting the  $^{163}\text{Ho}$  EC decay.
- Thanks to its low Q-value this isotope seems to be a good candidate for such a measurement.
- Some experimental challenges:
  - Embedding high activity into detectors arrays, managing possible pile-up issues;
  - Design and produce detectors arrays with suitable resolutions;
  - Implement a multiplexing readout scheme.
  - Optimization studies are ongoing.
- We expect to have the first data from implanted detectors by fall 2021.

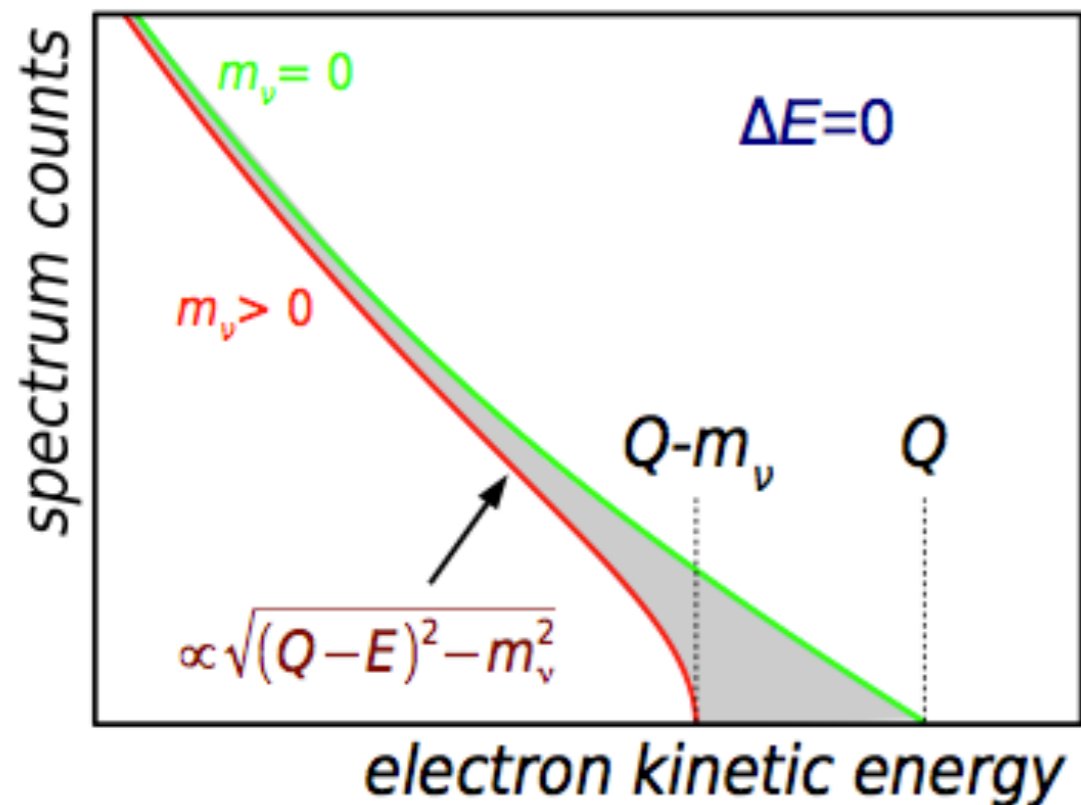


Back up slides

# Direct $\nu$ mass measurement

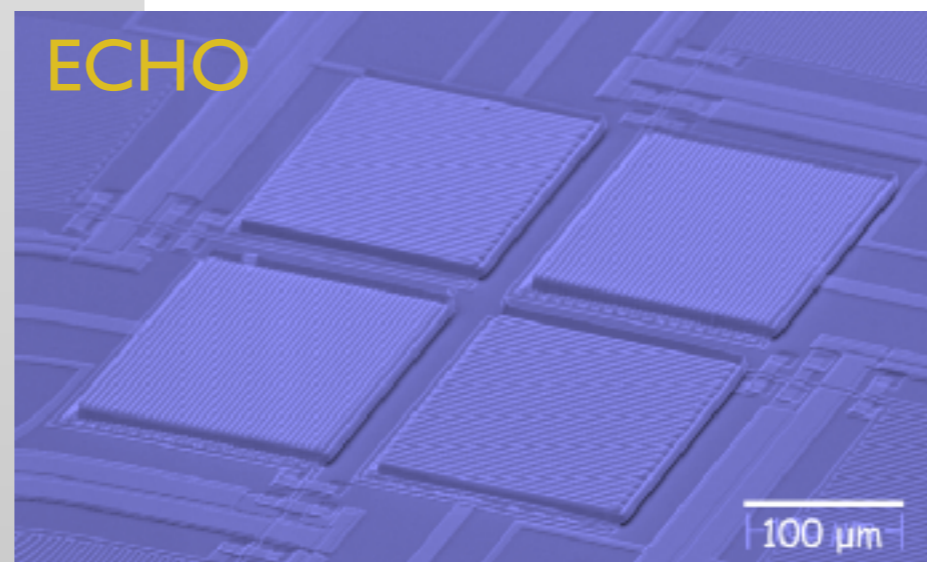
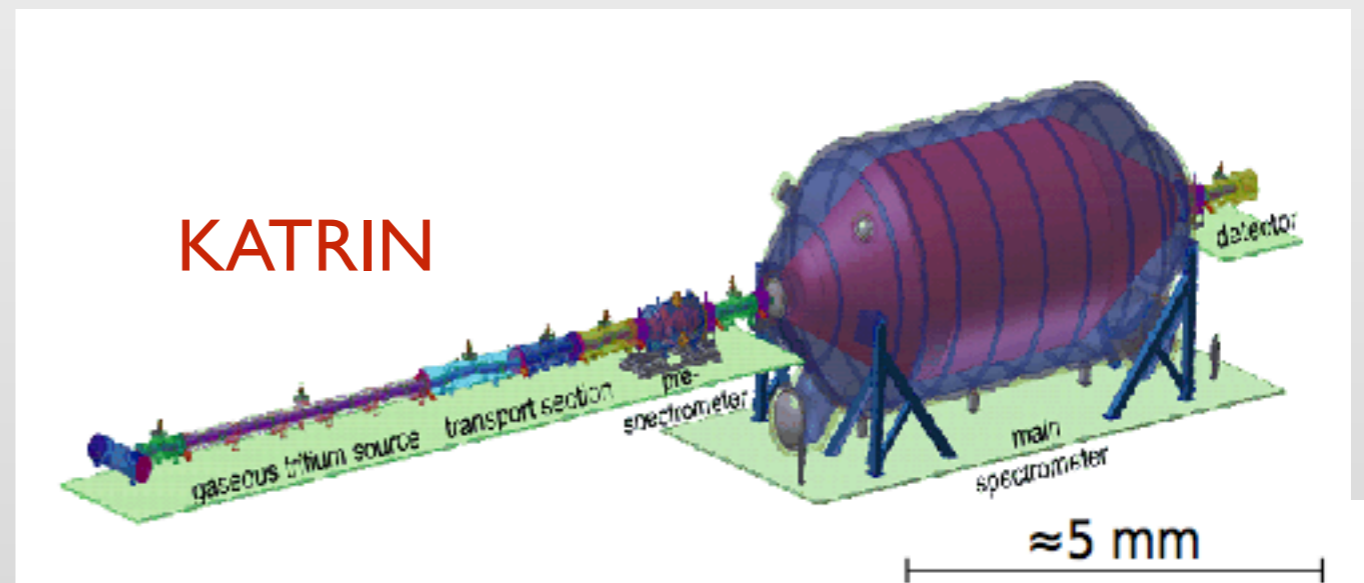
Kinematics of weak decay with  $\nu$  emission:

- low Q nuclear  $\beta$  decays ( $^3\text{H}$ ,  $^{187}\text{Re}$ ,  $^{163}\text{Ho}$ ...)
- model independent: only E, p conservation
- **$\nu$  mass appears as a distortion in the Kurie plot**



2 different approaches:

- **spectrometry**: source placed outside the detector (**KATRIN** approach)
- **calorimetry**: source embedded inside the detector (**ECHO**, **MARE**, HOLMES approach)  $\Rightarrow$  low T  $\mu$ -calorimeters



# Spectrometry vs calorimetry

---

General requirements for a  $\nu$  mass experiment:

- High statistics near the end point
  - low Q-value (stat  $\sim 1/Q^3$ )
  - high activity/efficiency of the source
- Energy reso order  $\sim eV$  or below (comparable with  $m_\nu$ )
- S/N ratio
- small systematic effects

Spectroscopy: source  $\not\subset$  detector

- high statistics
- high energy resolution (below eV)
- systematics due to the source (energy loss)
- systematics due to decay to excited states
- background

Calorimetry: source  $\subset$  detector

- no backscattering
- no energy loss in source
- no solid state excitation
- no atomic/molecular final state effects
- good energy resolution ( $\sim eV$ )
- limited statistics
- systematics due to pile-up
- background

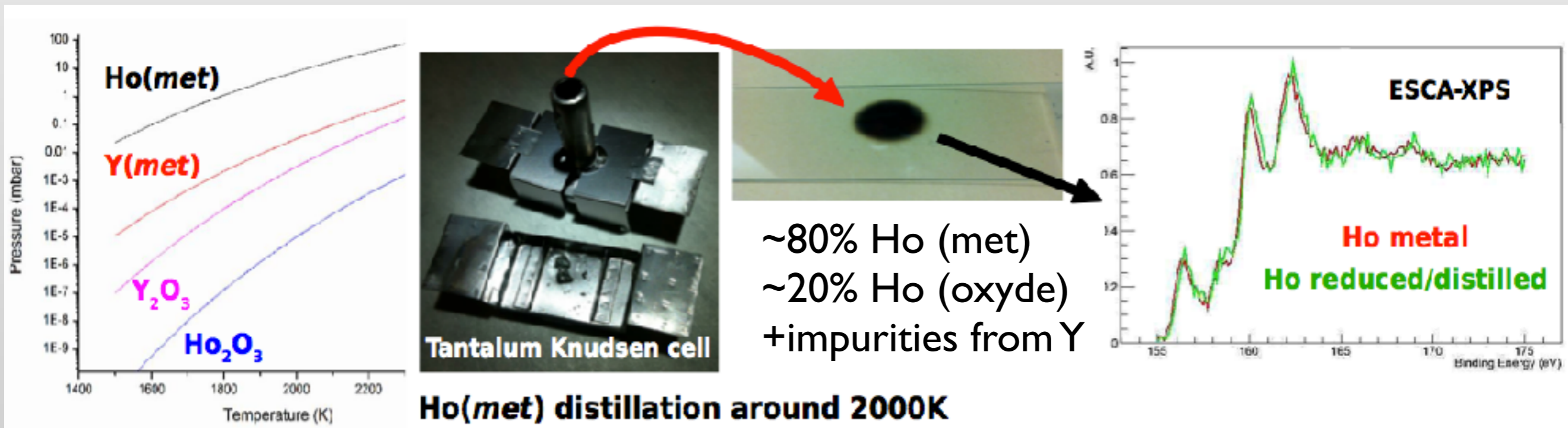
# Ho production and purification

$^{163}\text{Ho}$  separation from Dy, Er and others...

- radiochemistry (before/after activation process)
- magnetic mass separation

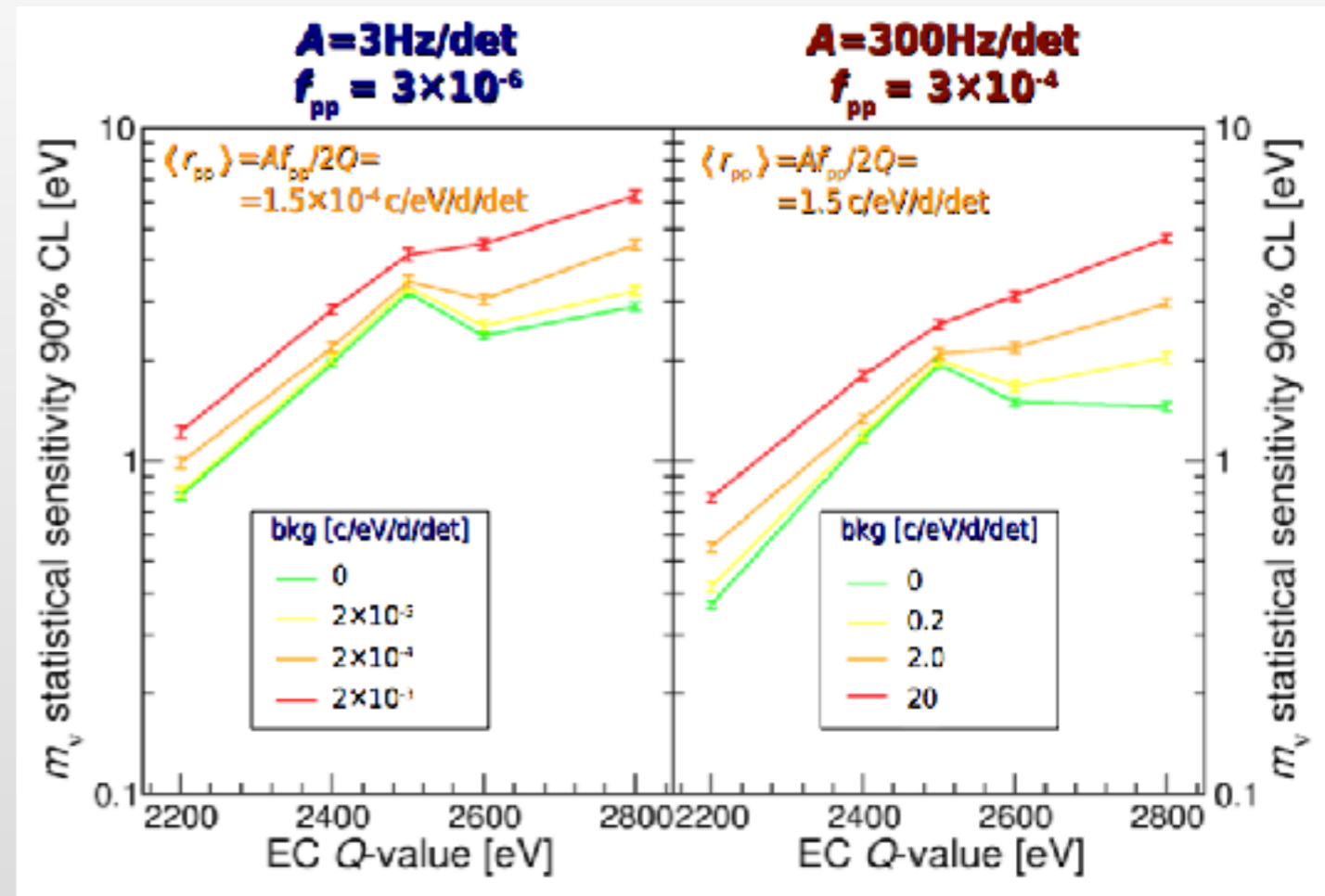
$\text{Ho}_2\text{O}_3$  thermoreduction in Knudsen cell provides a metallic sample for the implantation:

- $\text{Ho}_2\text{O}_3 + \text{Y}(\text{met}) \rightarrow \text{Ho}(\text{met}) + \text{Y}_2\text{O}_3 @ 2000\text{K}$
- First test already performed in Genova



# Source of background

- Environmental  $\gamma$  radiation
  - Compton interactions, photoelectric interactions with p.e. escape
  - Fluorescent X-rays and X-ray escape line
- Cosmic rays
  - GEANT4 simulation for CR at sea level (only muons)
  - Au pixel  $200 \times 200 \times 2 \mu\text{m}^3 \rightarrow \text{bkg} \sim 5 \times 10^{-5} \text{ c/eV/day/det}$  (0 - 4 keV)



- Internal radionuclides
  - $^{166\text{m}}\text{Ho}$  ( $\beta^-$ ,  $\tau_{1/2} = 1200 \text{ y}$ , produced along with  $^{163}\text{Ho}$ )
  - Au pixel  $200 \times 200 \times 2 \mu\text{m}^3 \rightarrow \text{bkg} \sim 0.5 \text{ c/eV/day/det/Bq}(^{166\text{m}}\text{Ho})$
  - $A(^{163}\text{Ho}) = 300\text{Bq/det}$  ( $\sim 6.5 \times 10^{13}$  nuclei/det)
  - if  $\text{bkg}(^{166\text{m}}\text{Ho}) < 0.1 \text{ c/eV/day/det}$ 
    - $\rightarrow A(^{163}\text{Ho})/A(^{166\text{m}}\text{Ho}) > 1500$
    - $\rightarrow N(^{163}\text{Ho})/N(^{166\text{m}}\text{Ho}) > 6000$