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- direct neutrino mass measurement
  - ▷ spectrometers vs. calorimeters
  - calorimeter statistical sensitivity
- ▷ <sup>187</sup>Re calorimetric experiment state-of-the-art
  - know systematics
- ▷ future of calorimetric experiments: the MARE project
- MARE project status
  - ▷ MARE-1
  - MARE R&D

Wisconsin University, Madison, July 27th, 2009

tool	measured quantity	present sensitivity	future sensitivity			
Cosmology CMB+LSS	$m_{\Sigma} \equiv \sum m_i$	0.7÷1 eV	0.05 eV	yes	large	
Neutrinoless Double Beta decay	$m_{\beta\beta} \equiv  \sum m_i U_{ei}^2 $	0.5 eV	0.05 eV	yes	yes	
Beta decay end-point	$m_{\beta} \equiv (\sum m_i^2  U_{ei} ^2)^{1/2}$	2 eV	0.2 eV	no	large	
		model dependency —				
		systematic uncertainties ——				

#### Neutrinos masses in single $\beta$ and $\beta\beta$ -0 $\nu$ decays



Fogli et al. hep-ph/0608060

# **Experimental approaches for direct measurements**

#### **Spectrometers: source ≠ detector**



#### **Calorimeters: source ⊆ detector**



#### $\beta$ calorimeter

ideally measures all the energy *E* released in the decay except for the  $\overline{v}_e$  energy:  $E = E_0 - E_v$ 

#### Spectrometers present results



# **Calorimetry of beta sources**

• calorimeters measure the entire spectrum at once  $\Rightarrow$  use low  $E_0 \beta$  decaying isotopes to achieve enough statistics near the end-point  $\Rightarrow$  best choice <sup>187</sup>Re:  $E_0 = 2.47$  keV  $\Rightarrow F(\delta E = 10 \text{ eV}) \sim (\delta E/E_0)^3 = 7 \times 10^{-8}$ 



#### Calorimetric experiment statistical sensitivity / 1



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#### Calorimetric experiment statistical sensitivity / 2

$$\frac{\text{signal}}{\text{bkg}} = \frac{\left| F_{\Delta E}(m_{\nu}) - F_{\Delta E}(0) \right| t_{M}}{\sqrt{F_{\Delta E}(0)t_{M} + F_{\Delta E}^{\rho\rho}t_{M}}} = \sqrt{t_{M}} \frac{A_{\beta}N_{det}\frac{\Delta E^{3}}{E_{0}^{3}}\frac{3m_{\nu}^{2}}{2\Delta E^{2}}}{\sqrt{A_{\beta}N_{det}\frac{\Delta E^{3}}{E_{0}^{3}} + 0.3\tau_{R}A_{\beta}^{2}N_{det}\frac{\Delta E}{E_{0}}}} = 1.7 \text{ for } 90\% \text{ C.L.}$$

$$\sum_{90} (m_{\nu}) \approx 1.13 \sqrt[4]{\frac{E_0^3 \Delta E}{A_\beta t_M N_{det}}} + 0.3 \frac{\tau_R E_0^5}{t_M N_{det} \Delta E}$$

Optimal energy interval  $\Delta E$  $\Delta E = max(0.56E_0\sqrt{\tau_R A_\beta}, \Delta E_{FWHM})$ 



#### Cryogenic detectors as calorimeters



•  $\Delta T(t) = E/C e^{-t/\tau}$  with  $\tau = C/G$  and G thermal conductance

#### **Resistive thermometers: thermistors**

- doped semiconductors at Metal-Insulator-Transition ( $N_c$ =3.74×10<sup>18</sup> cm<sup>-3</sup> for Si:P)
- at  $T \ll 10K \rightarrow$  phonon assisted variable range hopping conduction (VRH)

$$\rho(T) = \rho_0 \exp(T_0/T)^{\gamma}$$

- $ightarrow T_0$  increases with decreasing net doping N
- ►  $T < 1 \text{ K} \Rightarrow \gamma = \frac{1}{2}$  (VRH with Coulomb Gap)



#### Thermal detectors for calorimetric experiments

 $\label{eq:started_st$ 

metallic rhenium single crystals

- ► superconductor with  $T_c = 1.6 K$
- NTD thermistors
- MANU experiment (Genova)
- dielectric rhenium compound (AgReO<sub>4</sub>) crystals
  - Silicon implanted thermistors
  - MIBETA experiment (Milano)

 $\rightarrow m_{\nu} < \approx 15 \text{ eV}$ 

#### Systematics summary: calorimeters vs. spectrometers

#### Calorimetry systematics

- detector response function (energy dependence, shape,...)
- energy dependent background
- pile-up effects
- condensed matter effects: BEFS
- <sup>187</sup>Re decay spectral shape
- ...?

#### Spectrometer systematics

- decays to excited final states
- energy losses in the source
- e<sup>-</sup> T<sub>2</sub> elastic scattering
- spectrometer stability (HV)
- source stability (density, potential, charging...)
- energy dependent background
- •...?

completely different systematics!

## Montecarlo simulations: statistics and systematics

generate many (500-1000) simulated experiments

 $\triangleright$  calculate total  $\beta$  spectrum

 $\triangleright \ S(E) = (N_{\rm ev} \ (N_{\beta}(E,0) + f_{\rm pp}N_{\beta}(E,0) \otimes N_{\beta}(E\,0)) + b(E)) \otimes g(E)$ 

- $N_{ev}$  total  $\beta$  statistics
- $N_{\beta}(E,0)$  normalized <sup>187</sup>Re spectrum for  $m_{\nu} = 0$
- $f_{pp}$  fraction of unresolved  $\beta$  pile-up events
- b(E) background (usually constant)
- g(E) detector energy resolution function (usually gaussian)
- $\triangleright$  generate spectra introducing Poisson fluctuations in S(E)

 $\triangleright$  fit the spectra with standard technique

▷ obtain 90% C.L.  $m_{\nu}$  sensitivity  $\sum_{90} (m_{\nu})$  from  $\sqrt{(1.7\sigma)}$  of  $m_{\nu}^2$  distribution

Montecarlo input parameters vs. real experiment parameters

$$\triangleright N_{\rm ev} = N_{\rm det} t_{\rm M} A_{\beta}$$

$$\triangleright f_{pp} \approx au_{R} A_{\beta} ( au_{R} pprox au_{rise})$$

- Assessing systematic uncertainties with Montecarlo simulations
  - generate simulated experimental spectra with systematic effect
  - analyze spectra without effect
  - $\triangleright$  obtain  $\sum_{90} (m_{\nu})$  and  $\Delta m_{\nu}^2$  as function of effect magnitude

#### Sub-eV m<sub>v</sub> statistical sensitivity



#### Sub-eV m<sub>v</sub> statistical sensitivity / 2



#### Sub-eV m<sub>v</sub> statistical sensitivity / 3



#### Effect of background on statistical sensitivity



#### MC analysis of systematics: large arrays



#### MC analysis of systematics: more effects...



- linear term in background
- linear deviation from quadratic beta spectrum

# **BEFS: Re vs. AgReO<sub>4</sub>**



#### **BEFS: Beta Environmental Fine Structure** Modulation of the electron emission probability due to the

atomic and molecular surrounding of the decaying nucleus: it is explained by the wave structure of the electron (analogous of EXAFS)



#### A. Nucciotti, Madison, July 27<sup>th</sup> 2009 20

#### **BEFS in MARE**



#### Statistics and systematics summary

#### exposure required for 0.2 eV *m*, sensitivity

$\boldsymbol{A}_{\beta}$	$ au_{R}$	$\Delta E$	N <sub>ev</sub>	exposure
[Hz]	<b>[μs]</b>	[eV]	[counts]	[det×year]
1	1	1	0.2 1014	<b>7.6</b> 10 <sup>5</sup>
10	1	1	0.7 10 <sup>14</sup>	<b>2.1</b> 10 <sup>5</sup>
10	3	3	1.3 10 <sup>14</sup>	<b>4.1</b> 10 <sup>5</sup>
10	5	5	1.9 10 <sup>14</sup>	<b>6.1</b> 10 <sup>5</sup>
10	10	10	3.3 10 <sup>14</sup>	<b>10.5</b> 10 <sup>5</sup>

source of uncertainty	quantity describing the uncertainty	<i>maximum</i> <i>uncertainty for</i> $\Delta m_{\nu}^2 < 0.01 \text{ eV}^2$
error on energy resolution $\Delta E$	$\sigma_{ m err}(\Delta E)/\Delta E$	0.02
error on single pixel energy calibration K	$\sigma$ (K)/K	0.0004
spread in energy resolution $\Delta E$ in the array	$\sigma_{ m spread}(\Delta E)/\Delta E$	0.1
underlying constant background	$N_{ m bkg}/N_{ m ev}$	10-8

# <sup>187</sup>Re calorimetric experiment statistical sensitivity

 $\Sigma(m_{\nu}) \approx$  20 eV

MIBETA detectors with 
$$\Delta E_{FWHM} = 30 \text{ eV}, \tau_R = 1.5 \text{ ms}$$
 $\triangleright$  for  $A_{\beta} = 0.15 \text{ decay/s} \rightarrow f_{pp} = 2 \times 10^{-4}$  $\triangleright$   $t_M = 3.6 \text{ y} \times \text{det} \rightarrow 1.6 \times 10^6 \text{ events}$  $\triangleright$   $\sum_{exp} (m_{\nu}) = 15 \text{ eV}$ 

10 -14



$$\Sigma(m_{\nu})$$
= 2 eV

b for 
$$A_{\beta} = 0.3$$
 decay/s →  $f_{pp} = 3 \times 10^{-5}$   
 $\sum_{MC} (m_{\nu}) = 2$  eV with 2×10<sup>10</sup> events  
 $\sum_{M} = 2000$  y×det



 $\sum(m_{\nu}) = 0.2 \text{ eV}$ 

• detectors with  $\Delta E_{\text{FWHM}} = 1 \text{ eV}, \tau_{\text{R}} = 1 \mu \text{s}$   $\triangleright$  for  $A_{\beta} = 1 \text{ decay/s} \rightarrow f_{\text{pp}} = 10^{-6}$   $\triangleright \sum_{\text{MC}} (m_{\nu}) = 0.2 \text{ eV}$  with  $\rightarrow 2.5 \times 10^{13}$  events  $\triangleright t_{\text{M}} = 8 \times 10^5 \text{ y} \times \text{det}$ 

# A project for a New Rhenium Experiment: MARE

goal: a sub-eV direct neutrino mass measurement complementary to the KATRIN experiment

#### MARE-1

▷ new experiments with large arrays using available technology and ready to start as soon as possible (i.e. 2008..2009)



#### MARE-2

▷ very large experiment with a  $m_v$  statistical sensitivity close to KATRIN but still improvable: 5 years from now for further detector R&D



#### **MARE Project: interested institutions**

MARE: Microcalorimeter Arrays for a Rhenium Experiment Università di Genova e INFN Sez. di Genova Goddard Space Flight Center, NASA, Maryland, USA Kirkhhof-Institute Physik, Universität Heidelberg, Germany Università dell'Insubria, Università di Milano-Bicocca e INFN Sez. di Milano-Bicocca NIST, Boulder, Colorado, USA ITC-irst, Trento e INFN Sez. di Padova PTB, Berlin, Germany University of Miami, Florida, USA Università di Roma "La Sapienza" e INFN Sez. di Roma1 SISSA, Trieste Wisconsin University, Madison, Wisconsin, USA GSI Darmstad, Caltech, CNRS Grenoble, ... funded R&D

http://crio.mib.infn.it/wig/silicini/proposal/

# MARE-1: TES vs. silicon implanted thermistors

aim: high statistics measurement with a ready-to-use technology
 few eV statistical sensitivity in few years

 $\triangleright$  investigate systematics in thermal calorimeters with  $10^9 \div 10^{10}$  events

cross-check spectrometer results

#### MARE-1 SEMICON (MIBETA2)

U. Milano-Bicocca / INFN Sez. Mi-Bicocca U. Insubria / INFN Sez. Mi-Bicocca ITC-Irst / INFN Sez. Padova U. Wisconsin, Madison NASA/Goddard

about 300 element arrays
 well known Si implanted thermistors
 AgReO<sub>4</sub> crystals

#### MARE-1 TES (MANU2) U. Genova / INFN Sez. Genova U. Miami, Florida PTB Berlin, Germany

about 300 element arrays
 newly developed
 transition edge sensors
 Re crystals



cross check
 common effort on systematics
 joint analysis to improve limit



#### MARE-1: MC simulations vs. formula



# MARE-1 SEMICON: the NASA/Goddard XRS2 array

# **6×6 array:** optimized for X-ray spectroscopy $\rightarrow$ ASTRO-E2 mission **detectors**: silicon implanted thermistor with HgTe absorber at T = 60 mK



#### **MARE-1 SEMICON**

# NASA/GSFC XRS2-2 arrays 6x6 pixels flat AgReO<sub>4</sub> single crystals m ≈0.5 mg detector R&D phase results best operating T ≈ 90mK ΔE ≈ 30 eV, τ<sub>R</sub>≈ 250 µs





# MARE-1 SEMICON: statistical sensitivity from MC



setup ready for 8 arrays

- 288 AgReO<sub>4</sub> crystals
- $^{\rm o}$  now starting with 2 arrays (72 ch.)
- gradual deployment

further detector optimization









calibration source pulling string



connection<sup>\*</sup> boxes







# MARE-1 TES: Superconducting transition edge sensors

- superconductor thin films used inside the phase transition at  $T_c$ 
  - ▶ pure superconductors: Ir ( $T_c$ = 112 mK), W ( $T_c$ = 15 mK), ...
  - ▶ metal-superconductor bilayers ⇒ tunable  $T_c$  (20÷200 mK) : Mo/Cu, Tl/Au, Ir/Au, ...
    - high sensitivity ( $A \approx 100$ )  $\Rightarrow$  high energy resolution
    - high electron-phonon coupling ⇒ high intrinsic speed
    - low impedance ⇒ SQUID read-out ⇒ multiplexing for large arrays



# MARE-1 TES: sensor development

• Pulsed Laser Deposition of thin films: pure Ir or Ir bilayers

0.5 mm

- detectors with metallic rhenium absorbers<sup>100</sup>
- 300 channel array
- detector R&D goal:
  - 1 mg Re crystals with:  $\Delta E = 5 \text{ eV}$ ,  $\tau_{R} = 10 \text{ }\mu\text{s}$
  - a further step towards MARE-2
- two read-out options
  - JFETs with cold impedance transformer
  - frequency multiplexed SQUIDs (FDM)
    - first 3x3 FDM chip is under test now



#### MARE-1 TES: statistical sensitivity



▶ about  $3 \times 10^{10}$  events in 3 years  $\Rightarrow m_{\nu} < 1.8 \text{ eV}$ 





# MMC – Magnetic Micro Calorimeters (Heidelberg)





sensor design optimization for MARE-2
 rhenium absorbers is in progress
 ▷ meander pick-up coils without external *B* field

# MMC: recent achievements (Heidelberg)



**MKIDs R&D for MARE-2** 







• exploit the temperature dependence of inductance in a superconducting film

- qp detectors suitable for large absorbers
- **fast** devices for high single pixel activity  $A_{\beta}$  and low pile-up  $f_{pp}$

# high energy resolution multiplexing for very large number of

#### Sensitivity

 $\Delta E = 5 eV$ 

- $t_{\rm M}$  = 36000 detectors x 3 years
- $A_{\beta} = 20 \text{ c/s/det}$

• 
$$\tau_{rise} = 1 \ \mu s \Rightarrow m_{v} < 0.2 \ eV$$

• 
$$\tau_{\rm rise} = 100 \ \mu s \Rightarrow m_{\rm v} < 0.4 \ {\rm eV}$$

Technique largely still to be proved!!

#### **Interested institutions**

- INFN Milano-Bicocca
- INFN Roma (exp. Rich: R&D for CMB)
- ITC-irst
- Caltech
- CNRS Grenoble

#### MARE-2: topics in single pixel design



#### MARE extensions: <sup>163</sup>Ho electron capture measurement



- calorimetric measurement of non-radiative Dy atomic de-excitations (Coster-Kronig, Auger...)
- fraction of events at end-point may be as high as for <sup>187</sup>Re: depends on  $Q_{FC}$  ( $\approx$ 2.5 keV)

 $\blacktriangleright Q_{\rm FC}?$ 

- fewer active nuclei are needed ( $\tau \approx 4000$  y)
  - can be implanted in any suitable absorber
  - first implantation tests at ISOLDE are encouraging
- new NASA/Goddard TES arrays ( $\Delta E = 2 \text{ eV}$ ) can be implanted with <sup>163</sup>Ho

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- $\circ$  thermal calorimetry of <sup>187</sup>Re decay can give sub-eV sensitivity on  $m_{\nu}$
- the MARE project is taking off
- MARE-1 intermediate scale experiments are starting
- R&D for MARE-2 large scale sub-eV experiment is starting
   MMC R&D is already in progress
  - array microfabrication ha made huge progresses
  - sensor performances are close to what is needed for MARE
  - new MUX schemes are being developed
  - ▷ New ideas are coming up (MKIDs, <sup>163</sup>Ho)



#### MARE and the cosmological relic neutrino background

# MARE-2: 50000 detectors, 20 mg each 650 g of <sup>187</sup>Re 4×10<sup>-8</sup> counts/year... 🛞

A. G. Cocco, G. Mangano and M. Messina, arXiv:hep-ph/0703075v2

- $\circ \Sigma_{90}(m_v) = 0.1 \text{eV} \text{ for } N_{ev} = 10^{15}; \Delta E = 3 \text{ eV}; f_{pp} = 10^{-5}$ 
  - $\triangleright \tau_{R} = 1 \ \mu s$  and 10 decays/s per detector
  - $\triangleright$  3.2×10<sup>6</sup> detector×year...  $\otimes$

#### **Recognized systematics in calorimeters**

- detector response function (energy dependence, shape,...)
  - ▷ important in  $AgReO_4(\rightarrow)$
- energy dependent background
  - $\triangleright$  study low energy environmental and material radioactivity (  $\rightarrow$  )
- condensed matter effects: BEFS
  - ▷ observed in Re and AgReO<sub>4</sub>: improve modeling (→)
- pile-up effects
  - $\,\triangleright\,$  under investigation with MC methods ( $\rightarrow$ )
- analysis artifacts
  - ▷ to be studied by MC methods
- ▼<sup>187</sup>Re decay spectral shape
  - $\triangleright$  improve Buhring parametrization
- Iong term metastable excited states
  - should be negligible
- electron surface escape
  - ▷ should be negligible
- ▼...?
  - more statistics

#### **Detector response function**



X-ray peaks have tails on low energy side

• 1~6 keV X-rays in AgReO<sub>4</sub> have an attenuation length  $\lambda$  < 2  $\mu$ m

- are the response functions for X-rays and for  $\beta$ s from <sup>187</sup>Re decay the same?
- need for a good phenomenological description of the X-ray peak shape

# **MIBETA: Measurement of response function (2004)**

- external X-rays probe only detector surface
- escape peaks allow internal calibration  $\triangleright \lambda(6 \text{ keV}) \approx 3 \mu \text{m}$  $\triangleright \lambda(70 \text{ keV}) \approx 400 \mu \text{m}$  in AgReO<sub>4</sub>
- escape peaks are broad because of natural widths of atomic transitions







the hidden background is a source of systematic uncertainties



Go underground?

# MIBETA: BEFS analysis (2005)

#### **BEFS: Beta Environmental Fine Structure**

Modulation of the electron emission probability due to the atomic and molecular surrounding of the decaying nucleus: it is explained by the wave structure of the electron (analogous of EXAFS)



#### BEFS experimental evidence in <sup>187</sup>Re $\beta$ decay

■ in AgReO<sub>4</sub> less pronounced than in metallic rhenium



BEFS is a possible source of systematic uncertainties in <sup>187</sup>Re neutrino mass experiments

- ⇒ EXAFS measurements
- $\Rightarrow$  better models

#### MC analysis of systematics: BEFS





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#### MC analysis of systematics: BEFS



#### A better statistical analysis: analytical vs. MC

#### Montecarlo analysis

- Many experiment MC simulation
  - $\triangleright$  90% C.L.  $m_{\nu}$  sensitivity from  $\sqrt{(1.64 \sigma)}$  of  $m_{\nu}^2$  distributions
  - ▷ useful for statistical sensitivity and systematic effects analysis

#### • Simulation inputs

- $\triangleright$   $N_{\rm ev} = N_{\rm det} \times t_{\rm M} \times A_{\beta}$  total number of events
  - $N_{det}$  number of detectors
  - $t_{\rm M}$  measuring time
  - $A_{\beta}$  <sup>187</sup>Re activity for single detector
- $ightarrow f_{\text{pile-up}} \approx \tau_{\text{R}} \times A_{\beta}$  unresolved pile-up event fraction
  - $\tau_{\rm R} \approx \tau_{\rm rise}$  time resolution for pile-up identification
- $\triangleright$  g(E): detector energy resolution function (usually gaussian)
  - $\Delta E$  FWHM detector energy resolution

Assessing systematic uncertainties with Montecarlo simulations
 ▷ generate simulated experimental spectra with systematic effect
 ▷ analyze spectra without effect
 ▷ obtain Δm<sup>2</sup><sub>y</sub><0 as function of effect magnitude</li>

#### Neutrinos masses in single *β* and *ββ*-0*ν* decays



# <sup>187</sup>Re calorimetric experiment statistical sensitivity



#### MANU experiment (1999)

- 1.6 mg metallic rhenium single crystal
- one detector only
- Ge-NTD thermistor
  - ▷ *△E*=96 eV FWHM
  - $\triangleright$  symmetric and without tails
- 0.5 years live-time
  - $\triangleright$  6.0imes10<sup>6 187</sup>Re decays above 420 eV
  - $m_{\nu}^2 = -462 + 579_{-679} eV^2$
  - $\, \triangleright \, m_{\nu} < 26 \, \mathrm{eV} \, (95 \, \% \, \mathrm{C.L.})$
- first observation of BEFS in <sup>187</sup>Re decay











# MIBETA experiment: 2002/03



2.35

2.40

C. Arnaboldi et al., Phys. Rev. Lett. 91 (2003) 161802 M. Sisti et al, NIM A 520 (2004) 125

2.50

2.45

energy [keV]

**5**9

2.55





De Rujula and Lusignoli, Phys. Lett. 118B (1982) 429